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FEFF8.10

The FEFF Project  
Department of Physics  
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version 8.10  
28 November, 2000

## Abstract

*Ab initio* self-consistent real space Green's function code for simultaneous multiple-scattering calculations of x-ray-absorption spectra and electronic structure. Output includes extended x-ray-absorption fine structure (EXAFS), full multiple scattering calculations of x-ray absorption near edge structure (XANES), and projected local densities of states (LDOS). Calculations of x-ray scattering amplitudes (Thomson and anomalous parts) and nonresonant x-ray emission, spin dependent calculations of x-ray magnetic circular dichroism (XMCD) and spin polarized x-ray absorption spectra (SPXAS and SPEXAFS) are also possible, but less automated.

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# Chapter 1

## Synopsis

FEFF8.10 calculates extended x-ray-absorption fine structure (EXAFS), x-ray-absorption near-edge structure (XANES) and local densities of state (LDOS), using an *ab initio* self-consistent real space Green's function (RSGF) approach for clusters of atoms ( $Z < 99$ ), including polarization dependence. Calculations are based on an all-electron, relativistic Green's function formalism. The method combines both full multiple scattering and a high-order path expansion based on the Rehr-Albers multiple scattering formalism. Calculation of the x-ray elastic scattering amplitude, nonresonant x-ray emission and x-ray natural circular dichroism, spin dependent calculations of x-ray magnetic circular dichroism (XMCD) and spin polarized x-ray absorption (SPXAS and SPEXAFS) are also possible, but much less automated.

For a quick start we suggest that users try a few of the examples in Section 4. For details on use of the code, examples and suggestions for calculation strategies, see Sections 2, 3, and 4. For details about the algorithms used see the discussion for the appropriate module in Section 2 and, for more detail, the published references listed in Appendix C.

FEFF is written in ANSI Fortran 77, with the non-standard extension of double precision complex (`complex*16`) variables. It requires at least 18 megabytes (MB) of available memory (RAM) to run. For XANES calculations, one generally needs more memory (about 60 MB of RAM for a cluster of 100 atoms, about 170 MB for a cluster of 200 atoms, and so on depending on computer type). See Appendix B for installation instructions.

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Please contact the authors concerning any problems with the code. See Appendix F for trouble-shooting hints and problem/bug reports.

FEFF is copyright protected software and users must obtain a license from the University of Washington Office of Technology Transfer for its use. See Appendix A for complete copyright notice and other details. Documentation and information on how to obtain the code can be found at the FEFF Project WWW URL:

<http://leonardo.phys.washington.edu/feff/>

or by e-mail to the FEFF Project Coordinator at [feff@phys.washington.edu](mailto:feff@phys.washington.edu).

Please cite FEFF and an appropriate FEFF reference if the code or its results are used in published work. See Appendix C for a list of appropriate citations.

The main published reference for FEFF8 is: *Real Space Multiple Scattering Calculation and Interpretation of X-ray Absorption Near Edge Structure*, A.L. Ankudinov, B. Ravel, J.J. Rehr, and S.D. Conradson Phys. Rev. **B58**, 7565 (1998).

The reference for the additional spectroscopies added in FEFF8.10 is *Theory of solid state contributions to the x-ray elastic scattering amplitude*, by A. L. Ankudinov, and J. J. Rehr, Phys. Rev. B, **62**, 2437 (2000).

A review covering all earlier FEFF codes is: *Theoretical approaches to x-ray absorption fine structure*, J.J. Rehr, R.C. Albers Rev. Mod. Phys. **72**, 621-654 (2000).

Several new features have been added to FEFF8 compared to FEFF7 and earlier versions of FEFF, largely by Alexei Ankudinov. These include an automated self-consistent potential algorithm and calculations of LDOS, the Fermi level, and charge transfer. For version FEFF8.10 he also added calculations of other spectroscopies including the x-ray elastic scattering amplitude, x-ray natural circular dichroism and nonresonant x-ray emission. Bruce Ravel developed the full multiple scattering (FMS) algorithms, which are necessary both for LDOS and for accurate XANES calculations. The format of the FEFF8 document was also developed primarily by B. Ravel. Anna Poiarkova contributed two new approaches for calculating multiple-scattering Debye-Waller factors. Matthew Neville added an improved, portable output structure to facilitate file transfer between machines and the interface to XAFS analysis codes. The high-order multiple-scattering routines, pathfinder, and standard input/output routines were developed by Steven Zabinsky for FEFF5 and 6, and are still in use. The Hedin-Lundqvist plasmon-pole self-energy and phase shift routines were developed for early versions of FEFF in collaboration with Jose Mustre de Leon, Dan Lu and R.C. Albers and are still part of FEFF8.

FEFF is compatible with a number of XAFS analysis packages. For details see the links at the FEFF Project home page <http://leonardo.phys.washington.edu/feff/>. An extensive course

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dealing with XAFS analysis using FEFF and the FEFFIT code of M. Newville has been developed by B. Ravel; for details see <http://leonardo.phys.washington.edu/ravel/course/Welcome.html> and the FEFFIT home page <http://cars9.uchicago.edu/newville/feffit/>.

The authors thank many users of experimental versions of FEFF8 for feedback and suggestions. The authors also thank R.C. Albers, K. Baberschke, F. Bridges, C. Brouder, G. Brown, S.D. Conradson, G. Hug, M. Jaouen, and E. Stern for useful comments.

font	denotes
SMALL CAPS	names of programs
typewriter font	contents of files
'quoted typewriter font'	file names
ROMAN CAPITALS	names of cards in the 'feff.inp' file
<i>slanted typewriter font</i>	commands executed at a command line

Table 1.1: Typographic conventions in this document

## Chapter 2

# Input File Controls

The main program `FEFF` reads only one file created directly by the user, the file `'feff.inp'` and executes the various program modules described below. An auxiliary program (`ATOMS`), developed by Bruce Ravel, is provided which can generate the `'feff.inp'` file from crystallographic input parameters. Information about `ATOMS` can be found on the WWW at

<http://feff.phys.washington.edu/~ravel/atoms/>

This section describes `'feff.inp'` and the commands that tell `FEFF` what to do. It may be helpful to look at one or more of the sample input files in Section 4 while reading this section. The input file for `FEFF8.10` is similar to `FEFF` version 5 through 7, except that several additional options have been added to permit self-consistent potential generation, full multiple scattering XANES calculations, polarization dependence, and others listed in Section 2.1.

The input file `'feff.inp'` follows the free format style introduced by S. Zabinsky in previous versions of `FEFF` and is mostly backward compatible with previous versions. It is a loosely formatted, line oriented text file. Each type of input read by the program is on a line which starts with a *card* or *keyword* and, in some cases, is followed by alpha-numeric data. All card arguments listed below inside square brackets are optional. The sequence of keyword cards is arbitrary. If any card or optional data is omitted, default values are used; an exception is the `POTENTIALS` card, which is always required. Alpha-numeric values are listed in free format, separated by blanks. Tab characters are *not allowed* (due to Fortran 77 portability constraints) and may cause confusing error messages. Any characters appearing after the card and any required or optional data on a given line are ignored by `FEFF` and can be used as end-of-line comments. All distances are in Ångstroms and energies in eV. Spaces between lines (empty lines) are ignored. Any line beginning with an asterisk (\*) is regarded as a comment and is also ignored.

After reading the `'feff.inp'` file, the calculations are done sequentially in six steps:

1. The scattering potentials are calculated using atomic overlap (Mattheiss) prescription or self-consistently using an automated SCF loop. Absolute energies are estimated. (module `POT`, potentials).
2. The scattering phase shifts, dipole matrix elements, x-ray cross-section and angular mo-



mentum projected density of states (LDOS) are calculated. (module **XSPH**, cross-section and phases).

3. Full multiple scattering XANES calculations are done for a specified cluster size. (module **FMS**, full multiple scattering).
4. The leading multiple scattering paths for the cluster are enumerated. (module **PATHS**).
5. The effective scattering amplitudes  $f_{\text{eff}}$  and other XAFS parameters are calculated for each scattering path. (module **GENFMT**, general-path F-matrix calculation).
6. The XAFS parameters from one or more paths are combined to calculate a total XAFS or XANES spectrum. (module **FF2CHI**, scattering amplitude to chi).

This section describes how to control each module using the ‘**feff.inp**’ file.

## 2.1 Complete list of FEFF8 control cards

The list of ‘**feff.inp**’ options falls into three categories, *standard* options frequently and easily used, *useful* options that are often used, and *advanced* options that are seldom necessary or helpful. Every card in ‘**feff.inp**’ file will influence the calculations, starting from some module. Thus for better understanding how each module can be affected by the input cards, we list three categories (standard, useful, and advanced) for each module separately.

### Module 0 ..... RDINP

**Purpose:** Read input data

**Standard cards:** ATOMS, CONTROL, PRINT, and TITLE

**Useful Cards:** END, NOGEOM, and RMULTIPLIER

**Advanced Cards:** CFAVERAGE and OVERLAP

### Module 1 ..... POT

**Purpose:** Calculate (self-consistent) scattering potentials and Fermi energy

**Standard cards:** POTENTIALS, AFOLP, S02

**Useful Cards:** EXCHANGE, NOHOLE, RGRID, and SCF

**Advanced Cards:** FOLP, INTERSTITIAL, ION, and SPIN

### Module 2 ..... XSPH

**Purpose:** Calculate cross-section, phase shifts and  $\ell$ DOS

**Standard cards:** EXAFS, XANES, EDGE and HOLE,

**Useful Cards:** ELLIPTICITY, POLARIZATION, and LDOS

**Advanced Cards:** RPHASES, DANES, FPRIME, XES, XNCD

<b>Module 3</b> .....	FMS
<b>Purpose:</b> Calculate full multiple scattering for XANES	
<b>Standard cards:</b> FMS	
<b>Useful Cards:</b> DEBYE	
<b>Advanced Cards:</b>	
<b>Module 4</b> .....	PATHS
<b>Purpose:</b> Path enumeration	
<b>Standard cards:</b> RPATH	
<b>Useful Cards:</b> NLEG	
<b>Advanced Cards:</b> PCRITERIA and SS	
<b>Module 5</b> .....	GENFMT
<b>Purpose:</b> Calculate scattering amplitudes and other XAFS parameters	
<b>Standard cards:</b>	
<b>Useful Cards:</b> CRITERIA	
<b>Advanced Cards:</b> IORDER and NSTAR	
<b>Module 6</b> .....	FF2CHI
<b>Purpose:</b> Calculate EXAFS or XANES spectra	
<b>Standard cards:</b> DEBYE	
<b>Useful Cards:</b> CORRECTIONS and SIG2	
<b>Advanced Cards:</b>	

These data types are listed below in the same order as in the table above. Each description is of this form:

• **CARD** argument list *type*  
The type is one of *Standard*, *Useful*, or *Advanced*. The argument list is a brief statement of the valid arguments to the card. The text description will explain the arguments and their uses more fully. Example uses of the card look like this:

```
* brief description of the example
CARD argument list
```

## 2.2 Main Control Cards

These cards in this section are not associated with any particular module, but are used throughout the FEFF calculation. The ATOMS card is used to specify the absorbing atom and its environment. (If atomic coordinates are not known, then the OVERLAP card can be used to construct approximate potentials). Without this structural information no calculations can be done. The CONTROL card is used to selectively run parts of FEFF. The PRINT card controls which output files are written by the modules.

### • ATOMS

*Standard*

Cartesian coordinates (in Ångstroms) and unique potential indices of each atom in the cluster are entered following the ATOMS card, one per line. See also the discussion of the POTENTIALS card. An auxiliary code, ATOMS written by Bruce Ravel, is supplied with FEFF to generate the ATOMS list from crystallographic data. See the document file for ATOMS for more information.

```

ATOMS
* x      y      z      ipot      SF6 molecule
  0.0    0.0    0.0     0         S K-shell hole

  1.56   0.00   0.00    1         F 1st shell atoms
  0.00   1.56   0.00    1
  0.00   0.00   1.56    1
 -1.56   0.00   0.00    1
  0.00  -1.56   0.00    1
  0.00   0.00  -1.56    1

```

### • CONTROL ipot ixsph ifms ipaths igenfmt iff2chi

*Standard*

The CONTROL card lets you run one or more of the modules separately. There is a switch for each module: 0 means not to run that module, 1 meaning to run it. You can do the whole run in sequence, one module at a time, but you *must* run all modules sequentially. *Do not* skip modules: CONTROL 1 1 1 0 0 1 is incorrect. The default is CONTROL 1 1 1 1 1 1, i.e. run all 6 modules.

```

* example 1
* calculate self consistent potentials, phase shifts and fms only
CONTROL 1 1 1 0 0 0 ipot ixsph ifms ipaths igenfmt iff2chi

* example 2
* run paths, genfmt and ff2chi; do not run pot, xsph, fms
* this run assumes previous modules have already been run and
* adds MS paths between rfms and rpath to the MS expansion

```

```
CONTROL 0 0 0 1 1 1   ipot  ixsph  ifms   ipaths  igenfmt  iff2chi
```

• **PRINT** ppot pxsph pfms ppaths pgenfmt pff2chi *Standard*

The PRINT card determines which output files are printed by each module. The default is print level 0 for each module. See Section 3 for details about the contents of these files.

```
* add crit.dat and feffNNNN.dat files to minimum output
PRINT 0 0 0 1 0 3
```

The print levels for each module are summarized in Table 2.1 on page 9.

• **TITLE** title\_line *Standard*

User supplied title lines. You may have up to 10 of these. Titles may have up to 75 characters. Leading blanks in the titles will be removed.

```
TITLE Andradite (Novak and Gibbs, Am.Mineral 56,791 1971)
TITLE K-shell 300K
```

• **END** *Useful*

The END card marks the end of portion of the ‘feff.inp’ file read by FEFF. All data following the END card is ignored. Without an END card, the entire input file is read.

```
END   ignore any lines in feff.inp that follow this card
```

• **RMULTIPLIER** rmult *Useful*

With RMULTIPLIER all atomic coordinates are multiplied by the supplied value. This is useful to adjust lattice spacing, for example, when fractional unit cell coordinates are used. By default, rmult=1.

```
*increase distances by 1%
RMULTIPLIER 1.01
```

• **CFAVERAGE** iphabs nabs rclabs *Advanced*

A “configuration” average over absorbers is done if the CFAVERAGE card is given. CFAVERAGE currently assumes phase transferability, which is usually good for EXAFS calculations, but may not be accurate for XANES. Note that the CFAVERAGE card is currently incompatible with the NOGEOM card and with the DEBYE card for options other than correlated Debye model (idwopt > 0).

iphabs potential index for the type of absorber to over which to make the configuration average (potential 0 is also allowed)

module	print levels
pot	0 write 'pot.bin' and 'logpot.dat' only 1 add 'misc.dat' 2 add 'pot.dat' 3 add 'fpf0.dat' 5 add 'atomNN.dat'
xsph	0 'phase.bin', 'rkk.bin' and 'xsect.bin' only 1 add 'psisqNN.dat' and 'axafs.dat' 2 add 'phase.dat' and 'phmin.dat' 3 add 'ratio.dat' for XMCD normalization
fms	0 'gtr00.dat'
paths	0 'path00.dat' only 1 add 'crit.dat' 2 add 'geom.dat' 3 add 'fbeta' files (plane wave $ f(\beta) $ approximations) 5 Write only 'crit.dat' and save 'geom.dat' and do not write 'path00.dat'. (This is useful when exploring the importance of paths for large runs.)
genfmt	0 'list.dat', all paths with importance greater than or equal to two thirds of the curved wave importance criterion written to 'feff.bin' 1 keep paths written to 'feff.bin'
ff2chi	0 'chi.dat' and 'xmu.dat' 1 add 'sig2.dat' with Debye–Waller factors; 2 add 'chiNNNN.dat' ( $\chi(k)$ from each path individually) This can quickly fill up your disk if you're doing a large run. 3 add 'feffNNNN.dat' (input files for Matt Newville's FEFFIT program), and do not add 'chiNNNN.dat' files.

Table 2.1: Print levels controlling output files from the modules.

**nabs**

the configuration average is made over the first **nabs** absorbers in the 'feff.inp' file of type **iphabs**. You don't have to have potential of index 0 in your input file when using the CFAVERAGE card, but you must have the same type of potential for **iph=0** and **iph=iphabs**. The configurational average is done over ALL atoms of type **iphabs**, if **nabs** is less or equal zero.

**rclabs**

radius to make small atom list from a bigger one allowed in 'feff.inp'. Currently the parameter controlling the maximum size of the list **natxx**, is 100,000 but can be increased. The pathfinder will choke on too big an atoms list. You must choose **rclabs** to have less than 1,000 atoms in small atom list. If your cluster is less 1000 atoms simply use **rclabs=0** or negative always to include all atoms.

Default values are **iphabs=0**, **nabs=1**, **rclabs=0** (where **rclabs = 0** means to consider an infinite cluster size).

```
*average over all atoms with iph=2 in feff.inp with less than 1000 atoms
CFAVERAGE 2 0 0
```

• **OVERLAP iph**

*Advanced*

The OVERLAP card can be used to construct approximate overlapped atom potentials in terms of approximate near neighbor coordinations, when atomic coordinates are not known or specified. If the atomic positions are listed following the ATOMS cards, the OVERLAP cards are not needed. FEFF8.10 will stop if both ATOMS and OVERLAP cards are used. The OVERLAP card contains the potential index of the atom being overlapped and is followed by a list specifying the potential index, number of atoms of a given type to be overlapped and their distance to the atom being overlapped. The examples below demonstrate the use of an OVERLAP list. This option can be useful for initial single scattering XAFS calculations in complex materials where very little is known about the structure.

A user should verify that the coordination chemistry built in using the OVERLAP cards is realistic. It is particularly important to specify all the nearest neighbors of a typical atom in the shell to be overlapped. The most important factor in determining the scattering amplitudes is the atomic number of the scatterer, but the coordination chemistry should be approximately correct to ensure good scattering potentials. Thus it is important to specify as accurately as possible the coordination environment of the scatterer. Note: If you use the OVERLAP card then you cannot use the FMS or SCF cards. Also the pathfinder won't be called, and you must explicitly specify single scattering paths using SS card, which is described in Section 2.6.

\* Example 1. Simple usage

\* Determine approximate overlap for central and 1st nn in Cu

```
OVERLAP 0          determine overlap for central atom of Cu
```

```
*iphovr  novr  rovr      * ipot, number in shell, distance
```

```

1      12      2.55266
OVERLAP 1      determine approximate overlap for 1st shell atoms
*iphovr novr rovr * ipot, number in shell, distance
0      12      2.55266

```

\* Example 2. More precise usage

\* Determine approximate overlap for 3rd shell atoms of Cu

```

OVERLAP 3
0 1 2.55266      ipot, number in shell, distance
1 4 2.55266
2 7 2.55266
2 6 3.61000
2 24 4.42133

```

## 2.3 POT: Scattering potentials

FEFF8 has the capability for either non-self consistent (as in FEFF7 and earlier versions) or self-consistent scattering potentials. The calculation begins with free atom potentials of each atomic type (i.e., as if the atoms were isolated in space) using a relativistic Dirac–Fock atom code. Non-self-consistent scattering potentials are then calculated by overlapping the free atom densities within the muffin tin approximation (Mattheiss prescription), solving the Poisson equation, and for the Coulomb potential, and then including the complex Hedin–Lundqvist/Quinn self energy for excited states. Non overlapping muffin-tin radii are determined automatically from calculated the Norman radii. Automatic overlapping of muffin tin spheres (see the AFOLP card) is done by default, since it typically leads to better results than non overlapping muffin-tin spheres. Optionally FEFF8.10 determines self-consistent potentials by successively calculating the electron density of states, electron density and Fermi level at each stage within a small cluster and then iterating, starting from the Mattheiss prescription for the initial iteration.

XAFS spectra are referenced to the threshold Fermi level. This quantity is best determined with the self-consistent field procedure (typically to within a fraction of an eV), or (less accurately but much faster) can be estimated from the electron gas result at the mean interstitial density within Mattheiss prescription, as in FEFF7. An absolute energy scale is obtained by an atomic calculation of the total energy of the system with and without the core-hole. Atomic configurations and core-hole lifetimes are built in, and mean free paths are determined from the imaginary part of the average interstitial potential, including self-energy and lifetime contributions.

The potential calculations need as input only the atomic number of the atoms, and, for the absorbing atom, the type of the core hole being considered. To do the overlapping of the unique potentials, the neighboring atoms must be identified, either by position (from a list of the cartesian coordinates of each atom) or by explicit overlapping instructions using the OVERLAP card described in Section 2.3.

To save time the code calculates the overlapped atom potential for each unique potential

only once, using as a sample geometry for an atom with a given unique potential index that for the atom that is closest to the absorbing atom. Thus it is essential that the neighborhood of each sample atom be appropriate.

• **AFOLP** folpx

*Standard*

This automatically overlaps all muffin-tins up to a specified maximum value (default folpx=1.15) to reduce the effects of potential discontinuities at the muffin-tins. Automatic overlapping is done by default and is useful in highly inhomogeneous materials. Typical values of the overlapping fraction should be between 1.0 and 1.3. See FOLP for a non-automated version. Automatic overlapping is done by default; to switch off overlapping, use 1.0 as the argument for AFOLP.

\* touching muffin-tins; do not use automatic overlapping  
 \* recommended for AXAFS calculation  
 AFOLP 1.0

• **EDGE** label s02

*Standard*

The EDGE card is equivalent to the HOLE card, but eliminates the need to look up the appropriate integer index. Simply use the hole label: N0 means no hole, K means *K*-shell, L1 means  $L_I$ , and so on. As with the HOLE card you may also use the integer index instead of the label. All comments for HOLE card are valid for EDGE card - see the description below. Thus if the entry for  $S_0^2$  is less than 0.1,  $S_0^2$  will be estimated from atomic overlap integrals.

\* L1-shell core hole, S02 = 0.9  
 EDGE L1 0.9

• **HOLE** ihole s02

*Standard*

The HOLE card includes the hole-code index and the amplitude reduction factor  $S_0^2$ . If the entry for  $S_0^2$  is less than 0.1, then  $S_0^2$  will be estimated from atomic overlap integrals. Experimental values of  $S_0^2$  are typically between 0.8 and 1.0. The defaults if the HOLE card is omitted are ihole=1 for the *K* shell and  $S_0^2=1$ . The hole codes are shown in Table 2.2.

FEFF is designed to calculate absorption only from completely filled shells. You can try to simulate absorption from valence electrons with FEFF, but you may get unreliable results. If you encounter difficulties, and need valence shell absorption, please contact the authors.

For ihole > 4, the core-hole lifetime parameter ( $\gamma_{ch}$ ) is not tabulated in FEFF and is set equal to 0.1 eV, since the final state losses are then dominated by the self-energy. Use the EXCHANGE card to make adjustments ( $\gamma_{ch} = 0.1 + 2 \cdot vi0$ ).

\* K-shell core hole, S02 estimated by atomic overlap integrals  
 HOLE 1 0.0



index	edge	index	edge	index	edge	index	edge
0	N0	7	M3	14	N5	21	O5
1	K	8	M4	15	N6	22	O6
2	L1	9	M5	16	N7	23	O7
3	L2	10	N1	17	O1	24	P1
4	L3	11	N2	18	O2	25	P2
5	M1	12	N3	19	O3	26	P3
6	M2	13	N4	20	O4		

Table 2.2: Available hole codes. The entries in the columns marked edge are written as they are recognized by the EDGE card. Index 0, N0, is the no hole option described in the NOHOLE card.

### • POTENTIALS

*Standard*

The POTENTIALS card is followed by a list which assigns a unique potential index to each distinguishable atom. The potential index `ipot` is the index of the potential to be used for the phase shift calculation. The list is of this form

```
*   ipot  Z   [tag  lmax1  lmax2  xnatph]
```

The required list entries are the unique potential index `ipot` and the atomic number `Z`. The `tag` is at most 6 characters and is used to identify the unique potential in the ‘`path00.dat`’ output file. The optional list entries `lmax1` and `lmax2` are used to limit the angular momentum bases of the self-consistent potentials (XSPH) and full multiple scattering calculations (FMS), respectively. If the values are not specified, or if a negative number (e.g., `lmax1 = -1`) is specified for either `lmax1` or `lmax2`, FEFF will automatically use a default based upon the atomic number of the species normally `lmax(atomic)`. The last optional entry `xnatph` can be used to specify the stoichiometric number of each unique potential in the unit cell of a crystalline material. This helps in the calculation of the Fermi level. For an infinite solid, `xnatph = 0.01` (default value) is a suitable value for the absorbing atom (`ipot=0`).

The absorbing atom must be given unique potential index 0. These unique potential indices are simply labels, so the order is not important, except that the absorbing atom is index 0, and you may not have missing indices (i.e., if you use index 3, you must also have defined unique potentials 1 and 2).

To save time the code calculates the overlapped atom potential for each unique potential only once, using as a sample geometry the first atom in the atom list with a given unique potential index. Thus it is essential that the neighborhood of that sample atom be representative. Failure to do so may cause the code to generate inaccurate potentials and phase shifts and poor XAS results.

It is often useful to assume that the potential for a given shell of atoms is the same as that of a previously calculated shell in order to save calculation time. For example, in Cu it is a good approximation to determine potentials only for the central atom and the first shell and to use the first shell potential (`ipot = 1`) for all higher shells. Such approximations should be

checked in each case, however.

```
* molecular SF6 Sulfur K edge, lmax1=default, lmax2=3 (spdf basis)
POTENTIALS
*   ipot      Z   tag  lmax1 lmax2  xnat
      0       16   S    -1     3     1
      1        9   F    -1     3     6
```

• **S02** s02

*Standard*

The S02 card specifies the amplitude reduction factor  $S_0^2$ . If the entry for  $S_0^2$  is less than 0.1, then the value of  $S_0^2$  is estimated from atomic overlap integrals. Experimental values of  $S_0^2$  are typically between 0.8 and 1.0. This estimate is accurate to about 10% .

Alternatively, you can specify the value of  $S_0^2$  in the HOLE or EDGE card; however, the meaning of the parameters in the ‘feff.inp’ file is more clear if you use the S02 card.

```
* let FEFF calculate approximate S02
S02   0.0
```

• **FOLP** ipot folp

*Useful*

The FOLP card sets a parameter which determines by what factor the muffin-tin radii are overlapped. We recommend that the AFOLP card be used in cases with severe anisotropy, and FOLP only be used for diagnostic purposes. Typically only values larger than 1 and less than 1.3 should be used. The AFOLP card will be ignored once FOLP is used for any potential type: touching muffin tin radius is used for each potential type, unless it is changed with FOLP.

```
* 10% increase of touching muffin tin radius for unique potential 0
* 20% decrease of touching muffin tin radius for unique potential 1
* a 20% decrease is often needed for hydrogen atoms
FOLP 0  1.1
FOLP 1  0.8
```

• **EXCHANGE** ixc vr0 vi0 [ixc0]

*Useful*

The EXCHANGE card specifies the energy dependent exchange correlation potential to be used for the fine structure and for the atomic background. *ixc* is an index specifying the potential model to use for the fine structure and the optional *ixc0* is the index of the model to use for the background function. The calculated potential can be corrected by adding a constant shift to the Fermi level given by *vr0* and to a pure imaginary “optical” potential (i.e., uniform decay) given by *vi0*. Typical errors in FEFF’s self-consistent Fermi level estimate are about 1 eV. (The CORRECTIONS card in Section 2.8 is similar but allows the user to make small changes in *vi0* and *vr0* after the rest of the calculation is completed, for example in a fitting process.) The Hedin–Lundqvist self-energy is used by default and appears to be

the best choice for most applications we have tested in detail. The partially nonlocal model (`ixc=5`) gives slightly better results in some cases, but has not been tested extensively and will be improved in the future.

Another useful exchange model is the Dirac-Hara exchange correlation potential with a specified imaginary potential `vi0`. This may be useful to correct the typical error in non-self-consistent estimates of the Fermi level of about +3 eV and to add final state and instrumental broadening.

Defaults if EXCHANGE card is omitted are: `ixc=0` (Hedin-Lundquist), `vr0=0.0`, `vi0=0.0`. For XANES, the ground state potential (`ixc0=0`) is used for the background function and for EXAFS the Hedin-Lundqvist (`ixc0=0`) is used.

Indices for the available exchange models:

- 0 Hedin-Lundqvist + constant imaginary part
- 1 Dirac-Hara + constant imaginary part
- 2 Ground state + constant imaginary part
- 3 Dirac-Hara + HL imag part + constant imaginary part
- 5 Partially nonlocal: Dirac-Fock for core + HL for valence electrons + constant imaginary part

```
*Hedin-Lundqvist -2 eV edge shift and 1 eV expt broadening
EXCHANGE 0 2. 1.
```

```
*Dirac-Hara exchange -3 eV edge shift and 5 eV broadening
EXCHANGE 1 3. 5.
```

#### • NOHOLE

*Useful*

This card roughly simulates the effect of complete core-hole screening. It is useful to test the final state rule for calculated XAS, and to compare to other calculations (such as band structure or other codes) that do not use a core hole. The code will use as the final state that specified by the HOLE card for the matrix element calculation — the NOHOLE card will cause FEFF to calculate potentials and phase shifts as if there is no core hole. For  $d$ DOS and  $L_{II}$  or  $L_{III}$  absorption calculations, for example, NOHOLE often gives better agreement for white line intensities. Conversely NOHOLE tends to give poor XANES intensities for K-shell absorption in insulators such as BN.

#### • RGRID `delta`

*Useful*

The radial grid used for the potential and phase shift calculation is

$$r(i) = \exp(-8.8 + (i - 1) \cdot \delta)$$

with `delta` = 0.05 by default. The default is sufficient for most cases. However, occasionally there are convergence problems in the atomic background at very high energies (e.g., the

background curves upward) and in the phase shifts for very large atoms. If such convergence problems are encountered we suggest reducing `delta` to 0.03 (or even 0.01). This will solve these problems at the cost of longer computation times (the time is proportional to  $1/\delta$ ). This option is also useful for testing and improving convergence of atomic background calculations.

\* use smaller radial grid step for accurate high energy calculations  
RGRID 0.03

• **SCF** `rfms1 [lfms1 nscmt ca nmix]` *Useful*  
This card controls FEFF's automated self-consistent potential calculations. Thus all fields except `rfms1` are optional and not necessary except when some problems occur during SCF procedure. If this card is not specified then all calculations are done with the non self-consistent (overlapped atomic) potential. By default `lfms1=0`, `nscmt=30` and `ca=0.2`.

#### `rfms1`

This specifies the radius of cluster for full multiple scattering during the self-consistency loop. Typically one needs only about 30 atoms within sphere specified by `rfms1`, e.g. it is typically necessary to include first two coordination shells. Usually this value is smaller than the value `rfms` used in the FMS card, but should be larger than the radius of the second coordination shell.

#### `lfms1`

The default value 0 is appropriate for all solids; in this case the sphere defined by `rfms1` is located on the atom for which the density of states is calculated. The value 1 is appropriate for molecular calculations and will probably save computation time, but may lead to inaccurate potentials for solids. When `lfms1 = 1` the center of the sphere is located on absorbing atom.

#### `nscmt`

This is the maximum number of iterations the potential will be recalculated. A value of 0 leads to non-self consistent potentials and Fermi energy estimates. A value of 1 also yields non-self consistent potentials but the Fermi energy is estimated more reliably from calculations of the LDOS. Otherwise, the value of `nscmt` sets an upper bound on the number of iterations in the self-consistency loop. Usually self-consistency is reached in about 10 iterations.

#### `ca`

The convergence accelerator factor. This is needed only for the first iteration, since FEFF uses the Broyden algorithm to reach self-consistency. A typical value is 0.2; however, you may want to try smaller values if there are problems with convergence. After a new density is calculated from new Fermi level, the density after the first iteration is  $\rho_{\text{next}} = ca * \rho_{\text{new}} + (1 - ca) * \rho_{\text{old}}$ . `ca = 1.0` is extremely unstable and should not be used.

#### `nmix`

The `nmix` specifies how many iteration to do with mixing algorithm, before starting

Broyden algorithm. The calculations of SCF in materials which contain f-elements may not converge. We encountered such problem for Pu. However, SCF procedure converged if we started Broyden algorithm after 10 iterations with mixing algorithm with  $ca=0.05$ .

```
* Automated FMS SCF potentials for solids from 3.1 Ang clusters
SCF 3.1
```

```
* Automated FMS SCF potentials for a molecule of radius 3.1 Angstroms
SCF 3.1 1
```

```
* To reach SCF for f-elements we sometimes had to use
SCF 3.7 0 30 0.05 10
```

• **INTERSTITIAL inters vtot**

*Advanced*

The construction of interstitial potential and density may be changed by using INTERSTITIAL card. `inters = ipot + 2*irav + 6*irmt`. `ipot=1` might be useful when only the surroundings of the absorbing atom are specified in 'feff.inp'. The extra parameters were introduced for test purposes: `irav` and `irmt` are described only for completeness and nonzero values are strongly not recommended.

**ipot**

defines how to find interstitial potential: `ipot=0` (default) the interstitial potential is found by averaging over the entire extended cluster in 'feff.inp'. `ipot = 1` the interstitial potential is found locally around absorbing atom.

**irav**

also changes how interstitial potential is found. 0 (default) equation for  $V_{int}$  is constructed at  $r_{av}=r_{nrm}$ , 1 - at  $r_{av}=(r_{mt}+r_{nrm})/2$ , 2 - at  $r_{av}=r_{mt}$

**irmt**

0 : Norman prescription for mt radii (default) 1 : Matching point prescription for mt radii (do not use)

**vtot**

is the volume per atom normalized by the volume  $ratmin^3$ . (The quantity  $vtot = (volumeperatom)/ratmin^3$ ), where  $ratmin$  is the shortest bond for the absorbing atom. This quantity defines total volume (needed to calculate interstitial density) of the extended cluster specified in 'feff.inp'. If  $vtot \leq 0$  then the total volume is calculated as a sum of Norman sphere volumes. Otherwise,  $totalvolume = nat*(vtot*ratmin^3)$ ; where  $nat$  is a number of atoms in extended cluster. Thus  $vtot=1.0$  is appropriate for cubic structures, such as NaCl. The INTERSTITIAL card may be useful for open systems (e.g. those which have ZnS structure).

```
* improve interstitial density for ZnS structures.
```

```
* vtot = (unit_cell_volume/number_of_atoms_in_unit_cell)/ratmin**3)=1.54
INTERSTITIAL 0 1.54
```

• **ION** ipot ionization

*Advance*

The ION card ionizes all atoms with unique potential index ipot. Negative values and non-integers are permitted, however ionicities larger than 2 and less than -1 often yield unphysical results. Our experience with charge transfers using the SCF card suggests values for ionization about 5–10 times smaller than the formal oxidation state. The ION card is usually not needed if the potential is self-consistent. However, it can be used to put some total charge on a finite cluster. To do this we suggest using the same ionicity for all atoms in cluster (i.e. total ionization divided by number of atoms). For diatomics like Br<sub>2</sub>, the fully relaxed configuration has a formal ionization of 1 on the scattering atom. Because of charge transfer, however, the actual degree of ionization is will be much smaller. In non-self-consistent calculations the default (non-ionized) scattering potentials are often superior to those empirically ionized, and the results should be checked both ways. The default if ION cards are omitted is that the atoms are not ionized.

```
* Simulates effective ionization for formal valence state +1
* ipot, ionization
ION 1 0.1
```

```
* To put charge +1 on GeCl4 molecule use equal ionicity
ION 0 0.2
ION 1 0.2
```

• **SPIN** ispin [x y z]

*Advanced*

This card is used to specify the type of spin-dependent calculation for spin along (x, y, z) direction. The complete description is given in Section G.1, when dealing with spin-dependent calculations.

```
* XMCD calculation for spin along z-axis
SPIN 1
```

## 2.4 XSPH: Cross-section and phase shifts

Relativistic dipole matrix elements (alpha form) are calculated using atomic core and normalized continuum wavefunctions. Polarization dependence is optionally incorporated in the dipole-operator. Scattering phase shifts are determined by matching at the muffin-tin radius. Additionally,  $\ell$ -projected density of states can be calculated in this module, but it is of limited quality due to finite cluster calculations and neglect of nonspherical corrections.

• **ELLIPTICITY** *e x y z*

*Useful*

This card is used with the POLARIZATION card (see below).

The ellipticity  $e$  is the ratio of amplitudes of electric field in the two orthogonal directions of elliptically polarized light. Only the absolute value of the ratio is important for nonmagnetic materials. The present code cannot distinguish left and right circular polarization. A zero value of the ellipticity corresponds to linear polarization, and unity to circular polarization. The default value is zero.

$x, y, z$  are coordinates of any nonzero vector in the direction of incident beam. This vector should be approximately normal to the polarization vector.

\* light propagates along the x-axis, most of the light is polarized

\* along y-axis, and only about 4% of photons polarized along z-axis

```
ELLIPTICITY 0.2 1.0 0.0 0.0
```

```
POLARIZATION      0.0 1.0 0.0
```

\* circular polarization about the z-axis

\* can be used to average over polarization vector in xy-plane

```
ELLIPTICITY 1.0 0.0 0.0 1.0
```

• **POLARIZATION** *x y z*

*Useful*

This card specifies the direction of the electric field of the incident beam or the main axis of the ellipse in the case of elliptical polarization.  $x, y, z$  are the coordinates of the nonzero polarization vector. The ELLIPTICITY card is not needed for linear polarization. If the POLARIZATION card is omitted, spherically averaged xafs will be calculated.

Note that polarization reduces the degeneracy of the paths, increasing the calculation time. Choosing polarization in the directions of symmetry axes will result in a faster calculation.

```
POLARIZATION 1.0 2.5 0.0
```

• **EXAFS** [*xkmax*]

*Standard*

The EXAFS card is used to change the maximum value of  $k$  for EXAFS calculations. Default value is  $20 \text{ \AA}^{-1}$ . Now code can calculate even to higher values, however user will be prompted to increase dimensions in 'dim.h' file and must recompile the code. For high  $k$  calculations it might be necessary to make smaller radial step using RGRID card.

```
EXAFS 25
```

• **LDOS** *emin emax eimag*

*Useful*

The angular momentum projected density of states is placed by default on a standard grid currently fixed at 84 points. *emin* and *emax* are the minimum and maximum energies of the  $\ell$ DOS calculation and *eimag* is the imaginary part of potential used in the calculations. This is

equivalent to Lorentzian broadening of the  $\ell$ DOS with half-width = `eimag`. If `eimag` is negative, the code automatically sets it to be 1/3 of the energy step. The output will be written again into ‘`rho1MN.dat`’ files. To obtain LDOS you must run the second module by setting the second CONTROL argument to 1. If 84 points are not enough, you can divide the energy range by 2 and run the code twice. The LDOS card is very useful when examining densities of states for interpreting XANES or when the self-consistency loop fails or gives very strange results. For crystals our LDOS will always be broadened due to the effect of finite cluster size. The DOS is typically accurate for energies below  $E_{max} \approx lmax2^2$  in Hartree (atomic) units, where `lmax2` is specified in POTENTIAL card.

```
*      emin emax eimag
LDOS  -20  20   0.2
```

• **XANES** [`xkmax xkstep estep`]

*Standard*

The XANES card is used when a calculation of the near edge structure including the atomic background and absolute energies are desired.

The XANES calculation is currently limited to the (extended) continuum spectrum beyond the Fermi level. Thus bound states are not generally included; however, in molecules weakly bound states that are below the vacuum but above the muffin-tin zero will show up as resonances. The absolute energies are based on atomic total energy calculations using the Dirac-Fock-Desclaux atom code. The accuracy of this approximation varies from a few eV at low  $Z$  to a few hundred eV for very large  $Z$ . All parameters are optional. Default: XANES not calculated unless XANES card is present.

The optional parameters are used to change the output energy mesh for the XANES calculation. `xkstep` specifies the size of the output  $k$  grid far from the edge. `xkmax` is the maximum  $k$  value of the XANES calculation. FMS calculations are not accurate beyond about  $k = 6$ ; for larger values of  $k$ , e.g.  $k = 20$  with the path expansion, FMS must be turned off. The grid at the edge will be regular in energy with a step size of `estep`. The default values are `xkstep = 0.07`, `xkmax = 8`, and `estep = \gamma_{ch}/4 + vi0/2`, where `vi0` is given by the EXCHANGE card described in Section 2.3.

```
* finer grid for XANES calculation
XANES  6.  .05 .3
```

• **DANES** [`xkmax xkstep estep`]

*Advanced*

Calculates x-ray elastic scattering amplitude  $f'$  instead of absorption  $f''$ , including solid state effects. DANES calculates the contribution from a specified edge for a grid, which is specified exactly as in XANES card. This card is still experimental.

• **FPRIME** `emin emax estep`

*Advanced*

Calculates scattering factor  $f'$  far from the edge (only atomic part). The energy grid is regular



in energy with estep between emin and emax. This is typically needed to find out contribution from other edges to the edge calculated with DANES card. Later it may be automated. total scattering amplitude  $f'(Q, E) = f_0(Q) + f'(E) + if''(E)$  In the dipole approximation  $f'$  and  $f''$  do not depend on  $Q$ , which does not hold with quadrupole transitions added. This is currently neglected and  $f'(E) = \text{DANES}(\text{edge}) + \text{FPRIME}(\text{all} - \text{other} - \text{edges}) + \text{total} - \text{energy} - \text{term}$  in `fpf0.dat`.  $f''$  is printed out by FPRIME in electronic units, and can be used to obtain total  $f'$ . The total energy correction to  $f'$  is given in first line of '`fpf0.dat`' in Cromer-Liberman, and the more accurate Kissel-Pratt form. See our paper on the x-ray elastic scattering amplitude for references and details.  $f_0(Q)$  is tabulated in `fpf0.dat` for several values of  $Q$ .

• **XES** `emin emax estep`

*Advanced*

Calculates nonresonant x-ray emission spectra (XES) for a specified grid. It may be compared to DOS below Fermi level.

• **XNCD** [`xkmax xkstep estep`]

*Advanced*

Calculates x-ray natural circular dichroism. XNCD calculates contribution from specified edge and grid, which is specified exactly as in XANES card. This card has been used to calculate XNCD in our paper on elastic scattering amplitude (see list of references). This card is still experimental.

• **RPHASES**

*Advanced*

If this card is present, real phase shifts rather than complex phase shifts will be used. The results of the calculation will not be accurate. This option is intended to allow users to obtain real scattering phase shifts for use with other programs, or for diagnostic purposes. The phase shifts can be written to output files '`phaseNN.dat`' using the PRINT card. If the RPHASES card is present, these will be the real phase shifts.

• **RSIGMA**

*Advanced*

If this card is present, the imaginary part of self-energy will be neglected. This can be useful for calculations in XANES region, since the imaginary part of Hedin-Lundqvist self-energy tends to overestimate losses in this region. It is also useful for generating real phase-shifts with FEFF.

## 2.5 FMS: Full multiple scattering

This module carries out a full multiple scattering XANES calculation for a cluster centered on the absorbing atom. Thus all multiple-scattering paths within this cluster are summed to infinite order. This is useful for XANES calculations, but usually cannot be used for EXAFS analysis. FMS loses accuracy beyond  $k = (l_{\text{max}} + 1)/r_{\text{mt}}$ , which is typically about  $4 \text{ \AA}^{-1}$  since the muffin-tin radius  $r_{\text{mt}}$  is typically about  $1 \text{ \AA}$ .

• **FMS** `rfms lfms2`

*Standard*

Compute full multiple scattering within a sphere of radius `rfms` centered on the absorbing atom. If you don't use FMS card, the multiple scattering path expansion is used.

`rfms` is the cluster radius used in all modules but POT. Specifically is used for in the LDOS, FMS, and as the lower limit of pathfinder calculations. Typically a converged XANES calculation requires about 50-150 atoms in a cluster. The FMS module sums all MS paths within the specified cluster. The number of atoms in this cluster is limited to 87 by default, but one can manually increase the dimension parameter `nclusx` in the FEFF8.10 source code 'dim.h' (this has to be changed globally, since 'dim.h' is not a separate file in FEFF8.10) to alter the maximum cluster size. If there are more than `nclusx` atoms within the specified cluster size, FEFF will reduce the cluster size and issue a warning.

For EXAFS analysis one typically calculates to  $k = 20$ , but FMS results are not accurate at high energies. Thus if you are running FEFF8.10 for EXAFS you should not use FMS and XANES cards. It is, however, desirable to calculate self-consistent potentials even for EXAFS calculations as in the example below:

```
*calculate EXAFS with SCF potentials and paths to R=6 angstroms
CONTROL 1 1 1 1 1 1
*FMS
SCF 3.1
RPATH 6.0
EXAFS
```

If the value of RPATH as described in Section 2.6 is greater than `rfms`, the pathfinder will look for paths which extend beyond the cluster used for the FMS and add them to the FMS calculation of the  $\ell$ DOS and XANES:

$$G_{\text{tot}} = G_{\text{fms}} + G_0 t_i G_0 + G_0 t_i G_0 t_j G_0 + \dots$$

where at least one atoms  $i$  in the path is outside the FMS cluster and the value of RPATH is the maximum half path length for LDOS, FMS and pathfinder modules.

The MS expansion sometimes does not converge well in the XANES energy region. Thus one should avoid adding paths for LDOS and XANES, and RPATH should be less than `rfms`. Adding single scattering path only (NLEG 2) usually works well to check the convergence of FMS. But adding double scattering (NLEG 3) often leads to very bad results in XANES. Thus RPATH is useful for EXAFS or for XANES only when the path expansion is stable.

The optional `lfms2` argument is a logical flag which defines how the FMS is done, similar to the flag `lfms1` in the SCF card. With the default value of 0 (appropriate for solids), the FMS is calculated for a cluster of size `rfms` around each representative unique potential. With `lfms=1` (appropriate for molecules), FMS is done only once for a cluster of size `rfms` around absorbing atom only. The proper use of this flag can lead to a considerable time savings.

For example, if you calculate FMS for a molecule smaller than 40 atoms, there is no need to invert  $\text{nph} + 1$  matrices, and `lfms1 = 1` will reduce time for calculations by factor  $(\text{nph} + 1)$ , where `nph` is a number of unique potentials listed in POTENTIAL card).

A typical use of the FMS card uses `lfms2 = 0`, for example,

```
FMS  6.0      ! for XANES and LDOS need about 100 atom cluster
RPATH 8.0     ! usually use rpath < rfms
NLEG  2       ! adds 2 leg paths between 6 and 8 angstroms
```

For molecules of less than 30 atoms of radius 4.0 Å we suggest using `lfms2 = lfms1 = 1`, as in

```
FMS  5.0 1
RPATH -1
```

• **DEBYE** temp thetad [idwopt]

*Useful*

See the full description in Section 2.8 of the effect of temperature on EXAFS.

For XANES calculations the Debye-Waller factors are less important and the temperature effect can be approximated by multiplying each free propagator  $G_0$  by  $\exp(-\sigma^2 k^2)$ . This leads to the correct DW factors for single scattering  $\exp(-2\sigma^2 k^2)$ . The DW factors for multiple scattering are not exact, but their contribution is reduced both by thermal factors and by the mean free path. Also if you are using the FMS module, then you can obtain only XANES at low energies and this approximate treatment is probably adequate.

## 2.6 PATHS: Path enumeration

The code uses a constructive algorithm with several path importance filters to explore all significant multiple-scattering paths in order of increasing path length. The paths are determined from the list of atomic coordinates in ‘`feff.inp`’. An efficient degeneracy checker is used to identify equivalent paths (based on similar geometry, path reversal symmetry, and space inversion symmetry). To avoid roundoff errors, the degeneracy checker is conservative and occasionally treats two degenerate paths as not degenerate. These errors occur in the third or fourth decimal place (less than 0.001 Å) but are fail safe; that is, no paths will be lost. Of course, all paths which are completely inside the FMS cluster are automatically excluded from paths list.

The criteria used in filtering are based on increasingly accurate estimates of each path’s amplitude. The earliest filters, the pathfinder heap and keep filters, are applied as the paths are being searched for. A plane wave filter based on the plane wave approximation (plus a curved wave correction for multiple-scattering paths) and accurate to about 30% is applied after the paths have been enumerated and sorted. Finally, an accurate curved wave filter is applied to all remaining paths.

• **NOGEOM**

*Useful*

When this card is present, the file ‘`geom.dat`’ will not be produced. This option is useful to take

advantage of the symmetry in a cluster to speed the path calculation by using the ‘geom.dat’ file produced by the auxiliary code ATOMS rather than letting FEFF produce ‘geom.dat’ from the atom list in ‘feff.inp’. CAUTION: *DO NOT* use this option with the POLARIZATION card, since you would have to redo ‘geom.dat’ yourself for every polarization direction and you cannot use ATOMS to do that.

• **PCRITERIA** keep-criterion heap-criterion *Advanced*

These criteria, like those described in the CRITERIA card, also limit the number of paths. However, they are applied in the pathfinder and eliminate unimportant paths while the pathfinder is doing its search. The pathfinder criteria (pcrit’s) do not know the degeneracy of a path and are therefore much less reliable than the curved wave and plane wave criteria in the CRITERIA card above. These path finder criteria (keep and heap) are turned off by default, and we recommend that they be used only with very large runs, and then with caution.

The keep-criterion looks at the approximate amplitude of the XAFS  $\chi$  in the plane wave approx) for the current path and compares it to a single scattering path of the same effective length. To set this value, consider the maximum degeneracy you expect and divide your plane wave criterion by this number. For example, in fcc Cu, typical degeneracies are 196 for paths with large  $r$ , and the minimum degeneracy is 6. So a keep criterion of 0.08% is appropriate for a pw criteria of 2.5%.

The heap-criterion filters paths as the pathfinder puts all paths into a heap (a partially ordered data structure), then removes them in order of increasing total path length. Each path that is removed from the heap is modified and then considered again as part of the search algorithm. The heap filter is used to decide if a path has enough amplitude in it to be worth further consideration. If we can eliminate a path at this point, entire trees of derivative paths can be neglected, leading to enormous time savings. This test does not come into play until paths with at least 4 legs are being considered, so single scattering and triangular (2 and 3 legged) paths will always pass this test. Because only a small part of a path is used for this criterion, it is difficult to predict what appropriate values will be. To use this (it is only necessary if your heap is filling up, and if limiting  $r$ path doesn’t help), study the results in ‘crit.dat’ from runs with shorter  $r$ path and experiment with the heap criterion accordingly. In the future, we hope to improve this filter.

Before using these criteria, study the output in the file ‘crit.dat’ (use print option 1 for paths, see Table 2.1), which has the values of critpw, keep factor and heap factor for all paths which pass the critpw filter.

Default: If this card is omitted, the keep and heap criteria are set to zero, that is, no filtering will be done at this step in the calculation.

```
* fcc Cu had degeneracies from 6 to 196, so correct for this by
* dividing pw-crit of 2.5% by 30 to get 0.08 for keep crit. Check this
* empirically by running with pcrits turned off and studying crit.dat.
* After studying crit.dat, choose 0.5 for heap crit.
```

```
PCRITERIA 0.08 0.5
```

• **RPATH** `rpath`

*Useful*

The RPATH card determines the maximum effective (half-path) distance, `rpath`, of a given path. RPATH is equivalent to the RMAX card in the FEFF7 code. We changed the name because it provides a clearer distinction between the max distance in the MS path expansion and that for FMS calculations. Typically `rpath` is needed for EXAFS calculations only to set limits on the number of calculated paths. Note that `rpath` is one-half of the total path length in multiple-scattering paths. Setting this too large can cause the heap in the pathfinder to fill up. Default is `rpath` = 2.2 times the nearest neighbor distance. Since the multiple scattering expansion is unstable close to the absorption edge, the path (MS) expansion should be used only for EXAFS calculations or for diagnosing the XANES or LDOS calculations. If you use FMS for XANES calculations, better results are obtained without the MS contribution. For EXAFS analysis this card is extremely useful, since `rpath` cuts off long paths which give contribution only at high R values in R-space.

```
* include MS paths with effective length up to 5.10 Ang
RPATH      5.10
```

• **SS** `index ipot deg rss`

*Advanced*

The SS card can *only* be used with OVERLAP cards when the atomic structure is unknown but one does know the distance and coordination numbers and wants to generate an approximate EXAFS contribution. Thus the pathfinder cannot be used in this case. Instead the user has to specify explicitly the single scattering paths and their degeneracy. OVERLAP cards must be used to construct the potentials for the use with SS card. The parameters are a shell index, which is a label used for 'feffNNNN.dat' file name, a unique potential index `ipot` identifying the unique potential of the scattering atom, the degeneracy (or multiplicity) of the single scattering path, and the distance to central atom `rss`.

This information is used to write the file 'path00.dat' and is not needed when ATOMS card is used. To generate SS paths with ATOMS use (NLEG 2) card.

```
* index ipot  deg  rss  generate single scattering results
SS  29      1    48  5.98      parameters for 19th shell of Cu
```

## 2.7 GENFMT: XAFS parameters

For each path the code calculates the effective scattering amplitude ( $f_{\text{eff}}$  from which FEFF gets its name, see Section 3.2.4) and the total scattering phase shift along with other XAFS parameters using the scattering matrix algorithm of Rehr and Albers. Once the scattering phase shifts and the paths are determined, no other input is necessary for this calculation.

• **CRITERIA** `critcw critpw`

*Useful*

Since the number of multiple scattering paths gets large very quickly, it is necessary to eliminate as many paths as possible. Fortunately, we have found that most multiple scattering paths have small amplitudes and can be neglected. Various cutoff criteria are used in FEFF8.10 to limit the number of paths to consider. These criteria are based on the importance of the path, defined as the integral over the full energy range of  $\chi(k) \cdot dk$ . Very close to the edge these cutoff criteria should be examined with care and in some cases reduced from the values used for EXAFS.

`critcw` is the cutoff for a full curved wave calculation. A typical curved wave calculation requires a complete spherical wave calculation, which typically takes seconds of CPU time per path. The default value of `critcw` is 4%, meaning that any path with mean amplitude exceeding 4% of largest path will be used in calculation of chi. The criterion `critcw` is used by GENFMT. Since the XAFS parameter calculation is already done, the savings is not in computer time, but in disk space and ease of analysis. The values of `critcw` for each path are written in the file 'list.dat' written by module GENFMT.

`critpw` is a plane-wave approximation to  $\chi$ . This is extremely fast to calculate, and is used in the pathfinder. The default value of `critpw` is 2.5, meaning that any path with mean amplitude 2.5% of largest path, including degeneracy factors, (in plane wave approximation) will be kept. Any path that does not meet this criterion will not be written to 'path00.dat', and there is no need to calculate the XAFS parameters for this path. The default for `critpw` is less than that for `critcw` since some paths are more important when the full curved wave calculation is done than they appear in the plane wave approximation. Since the plane wave estimate is extremely fast, use this to filter out as many paths as you can. The file 'crit.dat' (written by the module PATHS) tells you `critpw` for each path that passes the criterion.

The method of calculation of these importance factors has been improved for FEFF8.10, so don't worry if the values for some paths have changed slightly from previous versions. (Default values `critcw`=4.% `critpw`=2.5%)

```
CRITERIA 6.0 3.0 * critcw 6%, critpw 3%
CRITERIA 0 0 * use all paths (cw and pw criteria turned off)
```

• **NLEG** `nleg`

*Useful*

The NLEG card limits the number of scattering paths to `nleg`. If `nleg` is set to 2, only single scattering paths are found. The default is `nleg` = 8.

```
* only single scattering paths (i.e. 2 legged paths)
NLEG 2
```

• **IORDER** `iord`

*Advanced*

Order of the approximation used in module GENFMT. FEFF uses order 2 by default which is correct to terms of order  $1/(pR)^2$ , and corresponds to 6x6 scattering matrices in the Rehr-Albers formalism. Single scattering is calculated exactly to this order. The 6x6 approximation

is accurate to within a few percent in every case we have tried (that is, higher order doesn't change the result more than a few percent). However  $M_{IV}$  shells and higher shells may require increased `iorder` for coupling the matrix elements. Changing the default values requires some familiarity with the Rehr–Albers paper and the structure of the module GENFMT. To do so, follow the instructions in the FEFF source code in subroutine `setlam`. The key `iord` is passed to `setlam` for processing. You may need to change the code parameter `lamtot` if you want to do higher order calculations. For details of the algorithm used by GENFMT, see the paper by J.J. Rehr and R.C. Albers (see the references in Appendix C). For the  $M_{IV}$  and higher edges, you may receive the error message like: **Lambda array overflowed**. In that case the calculations should be repeated with `IORDER -70202` (10x10 matrices).

```
* change iorder for M4 calculations
IORDER -70202
```

### • NSTAR

*Advanced*

When this card is present, GENFMT will write the file 'nstar.dat' with the effective coordination number  $N^*$  which is the coordination number weighted by  $\cos^2(\theta)$  to correct for polarization dependence in SEXAFS calculations. This is accurate only for the core hole in s-states (e.g. K, L<sub>1</sub> edges), but can be used as a guide for other edges.

## 2.8 FF2CHI: XAFS spectrum

The module FF2CHI constructs the XAS spectrum  $\chi(k)$  or  $\mu$  using the XAFS parameters described in Section 3.2.4 from one or more paths, including any FMS contributions. Single and multiple scattering Debye–Waller factors are calculated using, for example, the correlated Debye model. Output from this module is the total XAFS spectrum and optionally, the contribution to the XAFS from each path individually. Numerous options for filtering, Debye–Waller factors, and other corrections are available.

### • DEBYE temperature Debye-temperature [idwopt]

*Standard*

The Debye card is used to calculate Debye–Waller factors for each path using the correlated Debye Model. The model is best suited for homogeneous systems, where it is quite accurate. CAUTION: in heterogeneous systems the model only gives approximate values which can easily be off by factors of two or more. Temperatures are in Kelvin. If this card is present, the correlated Debye model Debye–Waller factors will be summed with the DW factors from the SIG2 card and from the 'list.dat' file, if present.

```
*Debye-Waller factors for Cu at 190K with correlated Debye Model
DEBYE 190 315
```

By default, `idwopt=0` specifies that the correlated Debye model is used to calculate EXAFS Debye–Waller factors. Two additional models for calculating DW factors are available

in FEFF8.10 based on the information about the harmonic force constants in the material. `idwopt=1` means the equation of motion (EM) method is used to get Debye–Waller factors and `idwopt=2` means the recursion method (RM) which is an improved correlated Einstein model. Both methods are faster than molecular dynamics simulations, and the recursion method is much faster than the equation of motion method. However, the equation of motion method leads to somewhat more accurate results than the recursion. These additional methods seem to be superior to correlated Debye model in cases with tetrahedral coordination, such as solid Ge or many biological materials. Both EM and RM methods need additional input (the force constants) and a complete description of both is given in Anna Poiarkova’s thesis (see the FEFF Project web site (<http://feff.phys.washington.edu>) and in the associated documentation.

```
* Calculate Debye-Waller factors for Cu at 190K with equation of motion
  DEBYE 190 0 1
```

• **CORRECTIONS** `real-energy-shift imaginary-energy-shift` *Useful*

The real energy shift moves  $E_0$  in the final  $\chi(k)$  and the imaginary energy shift adds broadening to the result. The real energy shift is useful to correct the error in FEFF’s Fermi level estimate and the imaginary part can be used to correct for experimental resolution or errors in the core-hole lifetime. This error in the Fermi level is typically about 1 eV with self-consistent calculations and about 3 eV with overlapped atom potentials. The imaginary energy is typically used to correct for instrument broadening or as a correction to the mean free path calculated by FEFF. This affects only the module FF2CHI, which combines the results in all of the ‘`feffNNNN.dat`’ files. This card is useful in fitting loops because you can simply make such energy corrections and see the results without redoing the entire XAFS parameter calculation. CAUTION: the results are not as accurate as those obtained with the EXCHANGE card. Both energies are in eV. (See also the EXCHANGE card in Section 2.3).

```
* Reduce E0 by 3.0 eV and add 1 eV of broadening (full width)
* This will only affect module 4, ff2chi
CORRECTIONS 3.0 1.0 real shift, imag shift
```

• **SIG2** `sig2` *Useful*

Specify a global Debye–Waller factor to be used or added to Debye–Waller calculations (see DEBYE card) for all paths. This value will be summed with the correlated Debye model value (if the DEBYE card is present) and any value added to ‘`list.dat`’. Units are  $\text{\AA}^2$ . This card can be used, for example to add Debye–Waller factors from structural disorder.

```
SIG2 0.001 add 0.001 globally to all DW factors
```



## Chapter 3

# Output Files

Other files required by the various modules are created by FEFF from ‘feff.inp’. Some of these other files may be edited by the user as a way to modify the input data to the modules, see Section 3.3. See the PRINT card in Section 2.2 to obtain various diagnostic files. Section 3.1 summarizes this structure, the rest of this section describes the structure in more detail.

### 3.1 Input and Output Files

**Module 0** ..... RDINP

**Purpose:** Read input data

**Input files:** ‘feff.inp’

**Output files:** ‘info.bin’, ‘geom.dat’, ‘potph.dat’, and ‘fms.dat’

**Other output:** ‘pathNN.dat’ (only if SS card is used)

**Description:** Reads the ‘feff.inp’ file, makes appropriate operations on the data, and writes resulting information into ‘info.bin’ and other files. The file ‘info.bin’ contains the data needed for all modules.

Note that ‘geom.dat’ is written when ‘feff.inp’ is read. If you are using a hand-made ‘geom.dat’ (possibly to take advantage of some symmetry), use the NOGEOM card to prevent your ‘geom.dat’ from being overwritten and save a copy.

**Module 1** ..... POT

**Purpose:** Calculate embedded atomic potentials for the photoelectron

**Input files:** ‘info.bin’, ‘potph.dat’

**Output files:** ‘pot.bin’

**Other output:** diagnostic files (see Table 2.1 on page 9)

**Description:** Reads ‘info.bin’ and ‘potph.bin’, both created by RDINP, and calculates potentials for the photoelectron which are written into ‘pot.bin’. Optionally, POT will write other diagnostic files with information about the potentials.

**Module 2** ..... XSPH

**Purpose:** Calculate cross-section and phase shifts

**Input files:** 'info.bin', 'potph.dat', and 'pot.bin'

**Output files:** 'phase.bin', 'xsect.bin', and 'rkk.bin'

**Other output:** diagnostic files (see Table 2.1 on page 9), 'axafs.dat', and 'rho1NN.dat' ( $\ell$ DOS)

**Description:** Reads 'info.bin', 'potph.dat', and 'pot.bin'. The file 'potph.dat' is deleted by FEFF after use. XSPH writes the binary file 'phase.bin', which contains the scattering phase shifts and other information needed by PATHS and GENFMT. Also, XSPH prints binary file 'rkk.bin', which contains dipole matrix elements, normalized to cross-section data in 'xsect.bin'. The last file is written only if XANES card is used. Optionally, XSPH will write other diagnostic files with information about the phase shift calculations.

**Module 3** ..... FMS

**Purpose:** Calculate full multiple scattering for XANES and  $\ell$ DOS

**Input files:** 'info.bin', 'fms.dat', 'phase.bin', 'rkk.bin', and 'geom.dat'

**Output files:** 'gtr00.bin'

**Other output:**

**Description:** Reads the 'info.bin', 'geom.dat', 'phase.bin' and 'fms.dat' files. Writes output into 'gtr00.bin' for the FF2CHI module, which contains the  $\chi(k)$  from FMS.

**Module 4** ..... PATHS

**Purpose:** Path enumeration

**Input files:** 'info.bin', 'phase.bin', and 'geom.dat'

**Output files:** 'path00.bin'

**Other output:** 'crit.dat'

**Description:** Reads 'geom.dat', which is created by FEFF from the information in 'feff.inp', and 'phase.bin'. PATHS writes 'path00.dat' for use by GENFMT and as a complete description of each path for use of the user. PATHS will optionally write other diagnostic files. The file 'crit.dat' is particularly useful when studying large numbers of paths. When studying large numbers of paths, this module will optionally write only 'crit.dat' and save 'geom.dat', not writing 'path00.dat'.

**Module 5** ..... GENFMT

**Purpose:** Calculate scattering amplitudes and other XAFS parameters

**Input files:** 'info.bin', 'phase.bin', 'rkk.bin', and 'path00.dat'

**Output files:** 'feff.bin', and 'list.dat'

**Other output:**

**Description:** GENFMT reads ‘phase.bin’, ‘rkk.bin’ and ‘path00.dat’, and writes a file ‘feff.bin’ which contains all the EXAFS information for the paths, and ‘list.dat’ which tells you some basic information about them. These files are the main output of FEFF. To read ‘feff.bin’ into your own program, use the subroutine feffdt as an example.

**Module 6** ..... FF2CHI

**Purpose:** Calculate XAFS spectrum

**Input files:** ‘info.bin’, ‘list.dat’, ‘feff.bin’, ‘gtr00.bin’, and ‘xsect.bin’

**Output files:** ‘chi.dat’ and ‘xmu.dat’

**Other output:** ‘chipNNNN.dat’ and ‘feffNNNN.dat’

**Description:** FF2CHI reads ‘list.dat’, ‘gtr00.bin’ and ‘feff.bin’, and writes ‘chi.dat’ with the total XAFS from the paths specified in ‘list.dat’. Additional instructions are passed to FF2CHI from ‘feff.inp’, so you can change S02, Debye temperature and some other parameters without re-doing the whole calculation. The file ‘list.dat’ can be edited by hand to change the paths being considered, and individual ‘chipNNNN.dat’ files with  $\chi(k)$  from each path are optionally written. If the XANES card is specified and ‘xsect.bin’ exists, FF2CHI will write the XANES data in ‘xmu.dat’. Various filtering options and corrections are possible at this point — see input cards above.

There is an internal limit on the number of paths (set to 1200) that will be read from ‘feff.bin’. This limit was chosen to handle any reasonable problem without using an excessive amount of memory. If you must use more paths, change the parameter npx in the FEFF source in subroutine ff2chi to whatever you need. It will need more memory. We have not had a case where the filter criteria were not able to solve the problem with fewer than 1200 paths.

## 3.2 Descriptions of output files

### 3.2.1 Intermediate output files

‘info.bin’

This binary file contains basic information from ‘feff.inp’ for all 6 modules.

‘pot.bin’

Charge density and potential (SCF or not) for all types of atoms. This file is used by XSPH module.

‘phase.bin’

This is a binary file with the scattering phase shifts for each unique potential. It is used by the FMS, pathfinder and GENFMT.

`'xsect.bin'`

Total atomic cross section for x-ray absorption.

`'rkk.bin'`

Relativistic dipole matrix elements, normalized to total cross section in `'xsect.bin'`.

`'rho1NN.dat'`

$\ell$ -projected density of states for the NN<sup>th</sup> potential index (see the LDOS card)

`'gtr00.dat'`

contains the results of FMS calculations. Used by FF2CHI to get total XAFS or XANES.

`'path00.dat'`

Written by the pathfinder, this is a description of all the paths that fit the criteria used by the pathfinder. It is used by GENFMT. The path descriptions include cartesian coordinates of atoms in the path, scattering angles, leg lengths and degeneracy. For details on editing this by hand, see Section 3.3. `'pathNN.dat'` files are created during the LDOS calculations for each type of potential, but they are deleted after use.

`'crit.dat'`

Values of the quantities tested against the various criteria in the pathfinder.

`'geom.dat'`

Written by feff for use with the pathfinder. It is an atom list created from the user-input list in `'feff.inp'`. For more information, see the NOGEOM card in Section 2.1.

`'list.dat'`

List of files to use to create `'chi.dat'`. Written by GENFMT when the xafs parameters are calculated and used by FF2CHI. This contains the curved wave importance ratios, which you may wish to study. For details on editing this by hand, see Section 3.3.

The curved wave importance ratios are the importance of a particular path relative to the shortest single scattering path.

### 3.2.2 Diagnostic files from XSPH

`'misc.dat'`

Header file for quick reference.

`'phaseNN.dat'`

Complex phase shifts for each shell.

`'phminNN.dat'`

Real part of phase shifts for  $\ell=0,1,2$  only. They are smaller versions of corresponding `'phaseNN.dat'`.

`'potNN.dat'`

Detailed atomic potentials and densities.

‘atomNN.dat’

Diagnostic information on Desclaux free atom NN.

### 3.2.3 Main output data

‘chi.dat’

Standard xafs data containing  $k$ ,  $\chi(k)$ ,  $|\chi(k)|$  relative to threshold ( $k = 0$ ). The header also contains enough information to specify what model was used to create this file.

‘xmu.dat’

The file ‘xmu.dat’ contains both XANES and XAFS data  $\mu$ ,  $\mu_0$ , and  $\tilde{\chi} = \chi\mu_0/\mu_0(\text{edge} + 50\text{eV})$  as functions of absolute energy  $E$ , relative energy  $E - E_f$  and wave number  $k$ .

‘feff.bin’

A binary file that contains all the information about the XAFS from all of the paths. This replaces the old ‘feffNNNN.dat’ files (which you can make using the PRINT card). If you want to use this file with your own analysis package, use the code in subroutine feffdt as an example of how to read it.

‘feffNNNN.dat’

You have to use the PRINT option to obtain these files. Effective scattering amplitude and phase shift data, with  $k$  referenced to threshold for shell nn:  $k$ ,  $\phi_c$ ,  $|F_{\text{eff}}|$ ,  $\phi_{\text{eff}}$ , the reduction factor,  $\lambda$ ,  $\text{Re}(p)$ .

If you need these, use the PRINT option for FF2CHI greater than 3, which will read ‘feff.bin’ and write the ‘feffNNNN.dat’ files in exactly the form you’re used to.

‘fpf0.dat’

Thompson scattering amplitude  $f_0(Q)$  and constant contribution to  $f'$  from total energy term.

‘ratio.dat’

Ratio  $\mu_0(E)$ ,  $\rho_0(E)$  and their ratio versus energy, for XMCD sum rules normalization.

### 3.2.4 Variables in the EXAFS formula

$k$  The wave number in units of  $\text{\AA}^{-1}$ .  $k = \sqrt{E - E_f}$  where  $E$  is energy and  $E_f$  is the Fermi level computed from electron gas theory at the average interstitial charge density.

$\chi(k)$

$$\chi(k) = S_0^2 \mathcal{R} \sum_{\text{shells}} \frac{N f_{\text{eff}}}{kR^2} \exp(-2r/\lambda) \sin(2kR + \phi_{\text{eff}} + \phi_c) \exp(-2k^2\sigma^2)$$

$\phi_c$  The total central atom phase shift,  $\phi_c = 2\delta_{\ell,c} - \ell\pi$

$f_{\text{eff}}$  The effective curved-wave backscattering amplitude in the EXAFS formula for each shell.

$\phi_{\text{eff}}$	The phase shift for each shell
$\mathcal{R}$	The total central atom extrinsic loss, $\mathcal{R} = \exp(-2 \text{Im}(\delta_c))$
$R$	The distance to central atom for each shell
$N$	The mean number atoms in each shell
$\sigma^2$	The mean square fluctuation in $R$ for each shell
$\lambda$	The mean free path in $\text{\AA}$ , $\lambda = -1/\text{Im}(p)$
$k_f$	The Fermi momentum at the average interstitial charge density
$p(r)$	The local momentum, $p^2(r) = k^2 + k_f^2(r) + \Sigma - \Sigma_f$
$\Sigma(E)$	The energy dependent self energy at energy, $\Sigma_f$ is the self energy at the Fermi energy.
$\mu(E)$	The total absorption cross-section
$\mu_0(E)$	The embedded atomic background absorption

### 3.3 Program control using intermediate output files

In addition to the CONTROL card and other options in ‘`feff.inp`’, some parameters in the files read by the various modules can be changed. For example, you can create your own paths by editing ‘`path00.dat`’ and explicit change Debye–Waller factors in the final result by editing ‘`list.dat`’.

Users may edit the some files as a quick and sometimes convenient way to prepare a given run. It is easiest to use an existing file as a template as the code which reads these files is fussy about the format of the files.

#### 3.3.1 Using ‘`path00.dat`’

You can modify a path, or even invent new ones, such as paths with more than the pathfinder maximum of 8 legs. For example, you could make a path to determine the effect of a focusing atom on a distant scatterer. Whatever index you enter for the path will be used in the filename given to the ‘`feffnnnn.dat`’ file. For example, for the choice of index 845, the EXAFS parameters will appear in ‘`feff0845.dat`’. A handy way to add a single scattering path of length  $R$  is to make a 2-leg path with the central atom at  $(0, 0, 0)$  and the scatterer at  $(R, 0, 0)$ .

GENFMT will need the positions, unique potentials, and character tags for each atom in the path. The angles and leg lengths are printed out for your information, and you can omit them when creating your own paths by hand. The label lines in the file are required (there is code that skips them, and if they’re missing, you’ll get wrong results).

**3.3.2 Using 'list.dat'**

This is the list of files that ff2chi uses to calculate chi. It includes the paths written by module GENFMT, curved wave importance factors, and user-defined Debye–Waller factors. If you want to set Debye–Waller factors for individual paths, you may edit this file to set them. FF2CHI will sum the Debye–Waller factors in this file with the correlated Debye model  $\sigma^2$  and the global  $\sigma^2$ , if present. You may also delete paths from this file if you want to combine some particular set of paths. (CAUTION: Save the original, or you'll have to re-run GENFMT!)

## Chapter 4

# Calculation Strategies and Examples

### 4.1 General comments

Self-consistent or overlapped atom potentials are needed for the calculation of the scattering phase shifts. Self-consistent calculations take more time and are often essential for XANES, especially for cases with significant charge transfer. Although the effect of self-consistency on EXAFS is small, such calculations give an accurate determination of  $E_0$ , thus eliminating an important parameter in EXAFS distance determinations.

Scattering phase shifts for each unique potential are necessary for FMS, PATHS and GENFMT. They are needed for the importance filters in PATHS and are the basis of the XAFS parameters calculation in GENFMT. This part of the calculation is relatively slow, so it is usually best to run it only once and use the results while studying the paths and XAFS.

To enumerate the necessary paths, the pathfinder module PATHS needs the atomic positions of any atoms from which scattering is expected. If the structure is completely unknown, only single-scattering paths can be created explicitly. Because the number of possible paths increases exponentially with total path length, one should start with a short total path length, examine the few paths (representing scattering from the nearest neighbors), and gradually increase the total path length, possibly studying the path importance coefficients and using the filters to limit the number of paths. This process is not automated, and if done carelessly can yield so many paths that no analysis will be possible.

Finally, use GENFMT to calculate the XAFS parameters, and FF2CHI to assemble the results into a chi curve. Here, the slow part is GENFMT and FF2CHI is very fast. Therefore, to explore parameters such as Debye–Waller factors, mean free path and energy zero shifts, various combinations of paths and coordination numbers, run only module FF2CHI using the results saved from GENFMT.

There are three ways to modify the Debye–Waller factor, all of which affect only the module FF2CHI. The DEBYE card calculates a Debye–Waller factor for each path. The SIG2 card adds a constant Debye–Waller factor to each path. And you can edit ‘list.dat’ to add a particular Debye–Waller factor to a particular path. These three Debye–Waller factors are



summed, so if the DEBYE and SIG2 cards are present, and if you have added a Debye–Waller factor to a particular path, the Debye–Waller factor used will be the sum of all three. See documentation below for details.

If your model changes significantly, the phase shifts (which are based in part on the structure of the material) should be recalculated. Any time the phase shifts change, the XAFS parameters will also have to be re-calculated. If the path filters have been used, the path list will also have to be recomputed.

## 4.2 EXAFS calculation

### 4.2.1 SF<sub>6</sub> Molecule

SF<sub>6</sub> Molecule. This is a simplest example of running FEFF to obtain EXAFS. This example is for fast, non self-consistent potential calculations. Only 2 input cards are necessary. Only 'chi.dat' file will be produced.

TITLE Molecular SF6

POTENTIALS

```
*   ipot   z   tag
      0    16   S      absorbing atom must be unique pot 0
      1     9   F
```

\* uncomment below for SCF calculations

\* SCF 4.0 1

ATOMS

```
*   x       y       z       ipot
      0       0       0         0       S absorber
     1.56     0       0         1       6 F backscatters
      0     1.56     0         1
      0       0     1.56     1
    -1.56     0       0         1
      0    -1.56     0         1
      0       0    -1.56     1
```

### 4.2.2 Solids

#### Cu metal

Cu, fcc metal, 4 shells. The list of atomic coordinates (ATOMS card) for crystals can be produced by program ATOMS. Thus instead of giving long atoms list, we present a short

'atoms.inp' file. For connection with EXAFS fitting programs see Section 3 and also the PRINT card on page 8.

```
TITLE Cu crystal, 4 shells
* Cu is fcc, lattice parameter a=3.61 (Kittel)

*Cu at 190K, Debye temp 315K (Ashcroft & Mermin)
DEBYE 190 315 0
```

```
POTENTIALS
  0 29 Cu0
  1 29 Cu
```

```
ATOMS
atoms list is generated using atoms.inp file below
-----
```

```
title Cu metal fcc a=3.6032
fcc ! shorthand for F M 3 M
rmax= 11.13 a=3.6032
out=feff.inp ! index=true
geom = true
atom
! At.type x y z
Cu 0.0 0.0 0.0
-----
```

### YBCO High-Tc superconductor

```
TITLE YBCO: Y Ba2 Cu3 07 Cu2 core hole
```

```
* FEFF7 input file can be used to run FEFF8
```

```
CONTROL 1 1 1 1
PRINT 0 0 0 0
```

```
RPATH 4.5
```

```
POTENTIALS
* ipot z tag
  0 29 Cu2
  1 8 0
  2 39 Y
  3 29 Cu1
  4 56 Ba
```

```

ATOMS
atoms list is generated by atoms.inp file
-----
title YBCO: Y Ba2 Cu3 O7 (1-2-3 structure)
space P M M M
rmax=5.2          a=3.823 b=3.886 c=11.681
core = Cu1
atom
! At.type  x      y      z      tag
  Y         0.5    0.5    0.5
  Ba        0.5    0.5    0.184
  Cu         0      0      0      Cu1
  Cu         0      0      0.356    Cu2
  O          0      0.5    0      O1
  O          0      0      0.158    O2
  O          0      0.5    0.379    O3
  O          0.5    0      0.377    O4
-----

```

### 4.2.3 Estimate of $S_0^2$

All above examples yield calculations for K edge (default). To do calculations for other edges use EDGE (or HOLE) cards. These cards will also yield an estimate of  $S_0^2$  from atomic overlap integral calculations if you set  $S02 = 0$  in those cards.

```
EDGE  L3  0.0
```

The result for  $S02$  is written in the header of 'chi.dat' or 'xmu.dat' files.  $S_0^2$  is a square of determinant of overlap integrals for core orbitals calculated with and without core hole. Core-valence separation can be changed by editing subroutine `getorb`, but it is currently set by default to the most chemically reasonable one.

### 4.2.4 Configuration averaging over absorbers

In amorphous materials or materials with distortions from regular crystals the absorbing atoms (with the same number in periodic table) may have different surroundings. Thus one may want to average the calculation over different types of sites for the same atom or even over all atoms in the 'feff.inp' file. This can be accomplished using CFAVERAGE card of Section 2.2.

```
*average over all atoms with iph=2 in feff.inp with less than 1000 atoms
CFAVERAGE 2 0 0
```

### 4.2.5 Adding self-consistency

Self-consistency is expected to be more important for the XANES calculations, but even for EXAFS one may want to have more reliable determination of Fermi level or to account for the charge transfers in order to do fits with single energy shift  $E_0$ . Our experience shows that reliable self-consistent results are best achieved by using both SCF and FMS cards.

```
* SCF for infinite systems (bulk or surfaces)
   SCF   3.8
```

The above example works for solids or large molecules, but for molecules with less than 30 atoms calculations can be done faster if you set `lfms1 = 1`,

```
* SCF for small molecules and clusters
   SCF  10.0  1
```

For details see SCF and FMS cards in Sections 2.3 and 2.5.

## 4.3 XANES calculations

### 4.3.1 Additional difficulties

XANES calculations are more challenging than EXAFS calculations. They usually take more time and require more experience from the user. Fortunately `FEFF8.10` automates many steps in the procedure and includes important full multiple scattering terms and self-consistency. This improves on the the high order path expansion approach to XANES in `FEFF7` which only allowed the maximum number of paths of amplitude larger than a plane wave criteria in `PCRITERIA` card. To account for the poor electron gas estimates of the Fermi level, `CORRECTIONS` cards were needed. Moreover, to obtain good results for the spectra, one had also to play with `AFOLP`, `EXCHANGE` and `ION` cards, which alter the way the scattering potential is constructed in somewhat uncontrolled ways.

One of the main advantages of the `SCF` card in `FEFF8.10` is that it can make a more reliable self-consistent estimate of Fermi level (the `CORRECTIONS` card can still be used, since the error in Fermi level position can still be about one eV or so) `FEFF8.10` also accounts for charge transfer. The `ION` card should be used only to specify the total charge of a cluster. `AFOLP` in general leads to better results for XANES and is done by default. With `FEFF7` we also had to use `EXCHANGE 5` model for Pu hydrates, but with the self-consistent `FEFF8.10` the standard `EXCHANGE 0` works well. Thus the use of self-consistency leads to closer results for different exchange correlation models.

The use of the high order MS path expansion and `PCRITERIA` can lead to unreliable XANES calculations when the MS series converges poorly (for example, near the Fermi level). Thus the inclusion of FMS capabilities in `FEFF8.10` is essential for calculations of LDOS and electronic densities and is often an improvement on path calculations of XANES. Actually we suggest that one use FMS exclusively for LDOS calculations and for testing convergence of the

path expansion. This will cost CPU time, but will lead to more reliable results. The FMS calculations for a cluster of 87 atoms typically take more time and memory than the other 5 modules. The results can be somewhat better with larger clusters, but typically one achieves convergence with about 50-200 atoms and calculational time scales as a third power of number of atoms in a cluster and quickly becomes prohibitive. Below we present several sample input files for XANES calculations.

### 4.3.2 GeCl<sub>4</sub> Molecule

This is a first molecule for which EXAFS was calculated by Hartree, Kronig and Peterson (1934), using the short range order theory.

```
TITLE   GeCl_4   r=2.09 /AA

HOLE 1   1.0
* NOHOLE is used to calculate absorption with ground state potential
* comment out below to see the effect of core-hole potential
NOHOLE

CONTROL  1  1  1  1  1  1

SCF      3.0  1
* uncomment below to calculate the density of states
*LDOS   -10  50  0.1
FMS      3.0  1
RPATH   1.0
XANES   8.0 0.05

AFOLP   1.30

POTENTIALS
*   ipot   z   label
      0   32   Ge -1 3
      1   17   Cl -1 3

ATOMS
*   x           y           z       ipot atom       distance
      0.0000      0.0000      0.0000      0   Ge
      1.2100      1.2100      1.2198      1   Cl
      1.2100     -1.2100     -1.2198      1   Cl
     -1.2100      1.2100     -1.2198      1   Cl
     -1.2100     -1.2100      1.2198      1   Cl

END
```

### 4.3.3 Solid

BN crystal has a zinc sulfide structure, and is a case in which a multiple scattering expansion does not converge near the Fermi level. Using the full multiple scattering approach leads to good agreement with experiment.

```
TITLE  BN cubic zinc sulfide structure
CONTROL  1 1 1 1 1 1
PRINT    5 0 0 0 0 0
```

```
SCF  3.1
HOLE 1  1.0      1=k edge, s0^2=1.0
EXCHANGE  0 0  1.0
LDOS -40  10  0.5
FMS  5.1
RPATH  1.0
XANES 4.0
```

```
INTERSTITIAL 0  1.54
```

#### POTENTIALS

```
*  ipot  z  label  lmax1  lmax2
      0   5   B      2      2
      1   7   N      2      2
      2   5   B      2      2
```

#### ATOMS

list generated by ATOMS program

```
-----
title BN (zincblende structure)
Space zns
a=3.615 rmax=8.0  core=B
atom
! At.type  x      y      z      tag
  B      0.0    0.0    0.0
  N      0.25   0.25   0.25
-----
```

### 4.3.4 Absolute cross-section

The absolute cross section can be obtained from the output in 'xmu.dat'. Look for the line:

```
xsedge+100, used to normalize mu          2.5908E-04
```

---

Multiplying the 4-th or 5-th column by this normalization value, yields the absolute cross section in  $\text{\AA}^2$ . ( $1 \text{\AA}^2 = 100 \text{ Mbarn}$ )

## Appendix A

# Copyright Information, Restrictions and License

### A.1 Restrictions and license information

FEFF is copyrighted software and a license from the University of Washington Office of Technology Transfer must be obtained for its use. This is necessary to protect the interests both of users and the University of Washington. Both academic/non-profit and commercial licenses are available — see Section A.2 to this document for details. New users should request the latest version of this code from the authors. The license form may be obtained from the FEFF WWW page,

<http://leonardo.phys.washington.edu/feff/>

or by writing or sending a FAX to

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Department of Physics, BOX 351560  
University of Washington  
Seattle, WA 98195-1560

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**A.2 ADDENDUM: Government Copyrights**

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## Appendix B

# Installation Instructions

The program FEFF8.10 is provided as a single source file ‘feff8.f’, which contains a main program and all necessary subroutines. Code dimensions are listed in each subroutine in a section labeled ‘dim.h’. [In the future, this may be moved to a separate ancillary file, but the present monolithic approach was adopted to ensure portability.] Simply compile (and link) on your system using a FORTRAN 77 (or FORTRAN 90) compiler (and your usual linker), e.g., for any UNIX system the command `f77 -o feff8 feff8.f` is usually sufficient. The executable file FEFF8.10 produced is all you will need. Typically FEFF8.10 requires about 64 MB of RAM to run for clusters of up to 87 atoms. To adjust the required memory to your computer, change the `nclusx` parameter in ‘dim.h’. The required memory scales approximately as `nclusx`<sup>2</sup>. For larger clusters, you may wish to reduce the maximum angular momentum `lx` to keep the total size reasonable, e.g., set `nclusx = 200, lx = 2`.

FEFF8.10 will also run on older ‘feff.inp’ files for FEFF7 (and also FEFF6) and will yield results for EXAFS and XANES comparable to older versions of FEFF. The first module of FEFF7 has now been split into 3 modules and therefore FEFF8.10 will use the first value in FEFF7 CONTROL and PRINT cards for the first 3 modules of FEFF8.10. But we suggest that users make use of the SCF, FMS and LDOS cards to take advantage of FEFF8.10’s new capabilities.

The FEFF code is written in ANSI Fortran 77, except that `complex*16` variables are required. Since data type `complex*16` is not part of the ANSI standard, minor compiler-dependent modification may be necessary. We have used the VAX extensions to ANSI Fortran 77 since they seem to be the most portable. The non-standard statements and intrinsic functions used are: `complex*16` variables and arrays; `dimag(arg)` returns a double precision imaginary part of its argument; `dcmplx(arg)` returns a `complex*16` version of its argument; `DBLE(arg)` returns a double precision real part of its argument; `sqrt`, `exp`, `abs` and other generic math functions are assumed to accept `complex*16` arguments and return double precision or `complex*16` results.

Files are opened with the `open` statement. As the ANSI definition gives some leeway in how the `open` statement interacts with the operating system, we have chosen file names and conventions that work on UNIX, VAX/VMS, IBM PCs (and clones) with MS FORTRAN,

CRAY, MAC's and CDC machines. It may be necessary to modify the open statements on other systems.

## B.1 UNIX

We have endeavored to make FEFF portable to most UNIX machines without any modification. If your machine does not reproduce the test output files 'xmu.dat' and/or 'chi.dat' to high accuracy, some changes may be needed. Please report any problems or warning messages to the authors, as this will help us achieve full portability. If changes are needed, keep in mind the following points:

Floating point calculations must be done to precision 64 bits. Thus the code uses double precision variables throughout, i.e., `real*8` for real numbers and `complex*16` for complex numbers. If your machine uses 8 bytes for single precision floating point numbers (for example, Crays and some CDC machines), use the compiler option to ignore double precision statements in the code. This option is available (in various forms) on all Cray and CDC machines that we have seen. On the CRAY YMP at Ohio State, the compiler option to ignore double precision in the source code is

```
cf77 -c -wf"-dp"
```

On SGI machines a significant increase in speed may be achieved with the following optimization flags

```
f77 -Ofast -LN0:opt=0 feff8.f -o feff8
```

On many unix systems it may be necessary to increase the memory stack size to allow calculations on large clusters. For example, for a 200 atom clusters, execute for following command before running FEFF8.10 if you use CSH or one of its derivatives

```
limit stacksize 170000
```

or this command if you use the Bourne shell or on of its derivatives:

```
ulimit -s 170000
```

If you use the GNU g77 compiler, a common compiler on LINUX, BSD, and other systems try for example:

```
g77 -O2 -ffast-math -m486 -Wall -g -fno-silent
```

The precise syntax can change, depending on the version of g77; in case of difficulty please contact the authors. Do not use the `-pedantic` flag, although a casual reading of the g77 document would suggest it is a good idea; `-pedantic` will not allow the use of double complex.

## B.2 MS-DOS, Win9x, and Win-NT

Because of the awkwardness of DOS, many users do not have FORTRAN compilers and many of those compilers are difficult to use with large codes. Thus we have made executable versions of FEFF for PCs available. You will need an 486 with a math coprocessor or pentium chip and 64 MB of RAM in addition to that needed for DOS and any -resident utilities. Further details are supplied with the executables.

If you prefer your own compiler (e.g., WATCOM or DEC/Compaq Visual Fortran), or are using an operating system other than DOS, compile the source code using your FORTRAN compiler and linker as you would for any other machine.

## B.3 Macintosh

FEFF is often difficult to compile on Macintosh machines, but Executable versions of FEFF for Macintosh computers are also available. Please contact us if you need a copy. For the G4 Macintosh computers the Absoft altivec compiler is recommended.

## B.4 VMS

To compile FEFF8.10 on VMS (6.0 to 6.2 versions) machine we had to increase virtual memory using SYSGEN from a SYSTEM account.

```
$mc sysgen
  sysgen> use current
  sysgen>show virtualpagecnt
  sysgen>set virtualpagecnt (value + 25%)
  sysgen>write current
  sysgen>exit
$reboot
```

If the code still does not compile one should further increase virtual memory. It may be impossible to compile the full code on old VAX stations. To run FEFF8.10 on VMS machines, you may need to increase quotas (*pgflquo=500,000*) for the users using AUTHORIZE from a SYSTEM account.

```
$run authorize
  authorize> modify username /pgflquo=500000
  authorize>exit
$reboot
```

## B.5 NeXT

Use the following options with Absoft's F77 for NeXT

---

```
f77 -N53 -f -s -O feff.f
```

*-f* makes F77 case insensitive. *-s* makes a code for units larger than 512K. *-N53* uses the 68030/68040 processors with 68881/2 math coprocessor. *-O* is an optimization flag.

---

## Appendix C

# References

Please cite at least one of the following articles if FEFF is used in published work.

### The primary reference for FEFF8

A.L. Ankudinov, B. Ravel, J.J. Rehr, and S.D. Conradson, *Real Space Multiple Scattering Calculation of XANES*, Phys. Rev. B **58**, 7565, (1998).

### A general review of modern EXAFS theory and previous versions of FEFF

J.J. Rehr, R.C. Albers, *Theoretical approaches to x-ray absorption fine structure*, Rev. Mod. Phys. **72**, 621-654 (2000).

### New features in FEFF8.1: $f'$ , DAFS, XNCD, etc.

A.L. Ankudinov, and J.J. Rehr, *Theory of solid state contributions to the x-ray elastic scattering amplitude*, Phys. Rev. B **62**, 2437 (2000).

FEFF7 A.L. Ankudinov and J.J. Rehr, *Relativistic Spin-dependent X-ray Absorption Theory*, Phys. Rev. B **56**, R1712 (1997).

A.L. Ankudinov, PhD Thesis, *Relativistic Spin-dependent X-ray Absorption Theory*, University of Washington, (1996). There you will find a review of x-ray absorption theory, a whole chapter of information about FEFF7 for expert users, examples of the code applications, and a full list of the FEFF7 program tree.

FEFF6 S.I. Zabinsky, J.J. Rehr, A. Ankudinov, R.C. Albers and M.J. Eller, *Multiple Scattering Calculations of X-ray Absorption Spectra*, Phys. Rev. B. **52**, 2995 (1995).

FEFF5 J.J. Rehr, S.I. Zabinsky and R.C. Albers, *High-order multiple scattering calculations of x-ray-absorption fine structure*, Phys. Rev. Let. **69**, 3397 (1992).

FEFF3 and FEFF4

J. Mustre de Leon, J.J. Rehr, S.I. Zabinsky, and R.C. Albers, *Ab initio curved-wave x-ray-absorption fine structure*, Phys. Rev. B **44**, 4146 (1991).

FEFF3 J.J. Rehr, J. Mustre de Leon, S.I. Zabinsky, and R.C. Albers, *Theoretical X-ray Absorption Fine Structure Standards*, J. Am. Chem. Soc. **113**, 5135 (1991).

Scattering theory in FEFF

J.J. Rehr and R.C. Albers, *Scattering-matrix formulation of curved-wave multiple-scattering theory: Application to x-ray-absorption fine structure*, Phys. Rev. B **41**, 8139 (1990).

Dirac-Fock atom code

A.L. Ankudinov, S.I. Zabinsky and J.J. Rehr, *Single configuration Dirac-Fock atom code*, Comp. Phys. Comm. **98**, 359 (1996).

## Appendix D

# Code Variables and Dimensions

The array names in FEFF are a bit cryptic due to the six character limit in standard FORTRAN — comments given in the list ‘`dim.h`’ (which is included in the header of each subroutine for code portability) and elsewhere in the source code explain what the names mean. If you need to run larger problems than the dimension statements in the code allow, you must change the dimensions globally in all the relevant parameter statements and recompile.

Thus you can adjust the main code dimensions in the ‘`dim.h`’ area of the headers to adjust the code for your computer or for a system of interest. If you want or need to modify the FEFF code for other reasons, please contact the authors. No part of FEFF can be included in other codes without a license or without permission from the authors. Our development version with all the comments (and subroutines in individual files!) is also available, although it makes use of some features that are not standard FORTRAN. We do not guarantee that any new version will be compatible at the code level with any changes you may make.



## Appendix E

# Changes From FEFF 7

FEFF8 potentials can now be calculated self-consistently (SCF card) which also gives a more accurate Fermi level position and accounts for charge transfer. Full multiple scattering capability was also added (FMS card). This is essential for SCF potentials, angular-momentum-projected density of states (LDOS also is a new card) and often XANES. All these new cards are not essential for EXAFS calculations, but the SCF potential can be used to reduce number of EXAFS fitting parameters by calculating the Fermi level and non-integral charge counts on each site. The possibility of calculating multiple-scattering Debye–Waller factors from force constants and/or dynamical matrices has also been added. The possibility of configurational averages (CFAVERAGE card) of EXAFS (or XANES) over different absorbers of the same type has been added. The CONTROL structure has been changed to accommodate the new cards, but backward compatibility has been maintained. Also several cards (e.g. EXAFS, XANES and POTENTIALS) now have additional optional fields.

The latest release FEFF8.10 is similar to the original release of FEFF8, but adds several new spectroscopies such as calculations of  $f'$ , DAFS, DANES, XES, and XNCD, which are also based on the same self-consistent, real-space multiple scattering formalism. Also an additional output file has been added 'ratio.dat' for the use in a new procedure for normalizing XAS sum rules.

FEFF8.10 also fixed a few bugs in the first release FEFF8.0. The most serious bug in FEFF8.10 was in polarization dependent calculations for initial states of non-*s*-character (e.g L2, L3 edges).

## Appendix F

# Trouble-Shooting FEFF Problems

FEFF has been extensively tested but occasionally new bugs show up. In an effort to maintain portable and trouble-free codes we take all bug reports seriously. Often we receive reports by users of older versions of FEFF of bugs that have been fixed in the latest releases. Other code failures can often be traced to input file errors, sometimes quite subtle.

Some known difficulties are:

- Non-physical widely spaced distributions of atoms. Symptoms of this common problem are very large muffin-tin radii (see the header of any `.dat` file) and possibly a failure of the phase-shift program to converge. This gives error message `hard test fails in fovrg`.
- An error in assigning potential indices; the first atom with a given potential index must have the geometry representative of this potential type. This is sometimes fixed by using a somewhat larger cluster; in fact it is usually desirable to have a larger cluster for potential construction than that used in the XAFS calculation due to errors in the potentials at surfaces. Unless the atom distribution is physically possible, you can expect the code to have problems.
- Hash collision in the pathfinder. This is now rare, but can usually be corrected simply by changing distances in the fourth decimal place.
- For the  $M_{IV}$  and higher edges you may receive the error message like: `Lambda array overflowed`. The calculations should be repeated with `IORDER -70202` card.

## Appendix G

# Advanced applications

### G.1 Spin dependent calculations

#### G.1.1 General description

These calculations are not currently automated. Code is not self-consistent with respect to spin variables and one has to specify the relative spin alignment and/or amplitudes by editing `ovrlp` subroutine. The default amplitudes are set in `getorb` subroutine. Ferromagnetic (all spins in the same direction) spin order is assumed for default, which is probably good enough to get corrections for the sum rules decribed below. Thus results in XANES may be questionable, but are often better than those obtained with other codes. Thus this code gave the best result to date for XMCD at Fe K edge, both in XANES and EXAFS, when compared to other calculations and experiment.

One can easily make mistakes in spin-dependent calculations with our code. It is probably best to ask Alexei Ankudinov for assistance. However, if you want to do them on your own, the spin-dependent calculations are described below. We plan to automate spin-dependent calculation in future versions of `FEFF`.

To accommodate spin dependent calculations, the `SPIN` card is added to `FEFF`. The values of the `ispin` argument correspond to:

<code>ispin</code>	meaning
-2	Calculate the spin-down SPXAS
2	Calculate the spin-up SPXAS
-1	Make the spin-down portion of XMCD calculations
1	Make the spin-up portion of XMCD calculations

Table G.1: Allowed values of the `ispin` argument of the `SPIN` card.

To get the XMCD signal you have to combine data from two ‘`xmu.dat`’ files. A simple program to do this, ‘`spin.f`’ is available on the `FEFF` web site, and also printed below.

These subroutines may need to be modified by the user:

#### getorb

The default values for spin amplitudes are set in this subroutine. The approximate atomic Hund's rule has been used to set values for d and f elements. Spin 1/2 is assigned for s, and p elements. This does not affect much XMCD calculations, however nonzero value is needed to get finite integration correction for these elements. You can reset the spin amplitudes by editing this subroutine and recompiling the code.

#### ovrlp

overlaps atomic densities and construct the total density magnetization, relative to the central atom. (SPIN card is used to define the sign of central atom density magnetization.) This subroutine requires user's attention, since the code is set for ferromagnets. There are commented out examples for antiferromagnet and ferrimagnet inside this subroutine. In later versions the relative spin orientation should be specified in 'feff.inp' (add 7-th column in POTENTIAL card). Currently it does not overlap magnetization, thus in interstitial region it is zero. It should be fine for antiferromagnets and f-element ferromagnets, but can lead to important corrections for d-element ferromagnets.

The spin-dependent potentials are calculated from the spin-dependent densities, using von Barth-Hedin results for the uniform electron gas.

We use the rough prescription to construct the spin-dependent muffin-tin potential. It should be fine for EXAFS where small details of the potential are irrelevant, but may be not good enough in the XANES region, where the self-consistent spin-dependent muffin-tin potential can lead to better results.

In order to use this spin dependent program you have to: 1) Check the construction of atomic density magnetization in the subroutine GETORB; 2) Check the construction of spin-dependent potential by OVRLP (examples for ferromagnets and antiferromagnets are there). 3) Be especially careful with antiferromagnets, since you may want to use the parity of iph to specify the relative directions of spin. 4) Now you can simply use the SPIN card to calculate SPXAS and XMCD. An additional simple program 'spin.f' is needed to take care of the different normalizations and give the finite results. If this experimental version will work, then later versions of FEFF will be automated to work with the SPIN card and the 7-th column of POTENTIALS to specify the relative spin directions on atoms.

An auxiliary program 'spin.f' can be used to get XMCD or SPXAS

```

implicit double precision (a-h,o-z)
c   This program read two xmu.dat files for spin -up and -down,
c   calculated with Feff8.01 for the SAME paths list.
c   spin-up file is fort.1, spin-down file is fort.2
c   Both have to be edited: All lines should be deleted except
c       1) line: xsedge+100, used to normalize mu           1.3953E-04
c           leave only on this line: 1.3953E-04
c       2) 6-column data lines

```

```

c   The output will be written in fort.3 in 6 columns
c   E+shift1 E(edge)+shift2 xk cmd_total cmd_background cmd_fs
c   where total = atomic background + fine structure

c   There are 3 possibilities
c   case 1) you want XMCD signal and used SPIN \pm 1
c   case 2) you want XMCD signal and used SPIN \pm 2, in order
c   to use nonrelativistic formula for XMCD
c   factor  $l_i/2j+1$  which was not convenient to do in a program
c   case 3) you want SPXAFS and used SPIN \pm 2
c   ENTER your case here (icase is positive integer only)
c   icase = 2

c   if icase=2 ENTER factor= $(-1)^{L+1/2-J} * L/(2J+1)$ 
c   where L,J are for your edge (ex. for L3 L=1 J=3/2, for L2 L=1 J=1/2)
c   for L3
c   factor = 0.25
c   for L2
c   factor = -0.5

c   ENTER the energy shift you want for columns 1 and 2 in xmu.dat
c   shift1 = 0
c   shift2 = 0

c   everything below is automated further
c   read (1,*,end=10) ap
c   read (2,*,end=10) am
c   xnorm = 0.5 *(ap+am)
c   read the data
3   read(1,*,end=10)  x1, x2, ek, y1, y2, y3
   read(2,*,end=10)  x1, x2, ek, z1, z2, z3
   if (icase.eq.1) then
c       no xafs in this case: xfs - atomic part of XMCD
       t1 = (y1*ap + z1*am)/xnorm
       t2 = (y2*ap + z2*am)/xnorm
       t3 = (y3*ap + z3*am) /xnorm

       elseif (icase.eq.2) then
       t1 = (y1*ap - z1*am)*factor /xnorm
       t2 = (y2*ap - z2*am)*factor /xnorm
       t3 = (y3*ap - z3*am)*factor /xnorm

       elseif (icase.eq.3) then
c       factor=0.5 always for SPXAFS

```

```

        t1 = (y1*ap - z1*am)/2.0/xnorm
        t2 = (y2*ap - z2*am)/2.0/xnorm
        t3 = (y3*ap - z3*am)/2.0/xnorm
c      you may want average total XAS as output in last column
c      t3 = (y1*ap + z1*am)/2.0/xnorm
      endif
      x1 =x1 + shift1
      x2 =x2 + shift2
      write(3,5)    x1, x2, ek, t1, t2, t3
5     format (6e13.5)
      goto 3
10    continue
      stop
      end

```

You really want to have the same paths used for spin up and down calculation, otherwise the difference between 2 calculations may be due to different paths used. Typically the paths list in 'path00.dat' should be generated by running the usual EXAFS calculations and comparing with experiment (to make sure that all important paths included). Then when running with SPIN turn off the pathfinder module using CONTROL card. This is probably the only place when you have to skip the pathfinder module. There is no rule without exception.

### G.1.2 XMCD

XMCD has recently become popular due, in part, to the utility of many sum rules in XANES region. The EXAFS region can be used to determine the position of spins relative to magnetic field. Consider for example, the Gd  $L_1$  edge.

```

TITLE   Gd  l1  hcp

HOLE 2   1.0      2=11 edge, s0^2=1.0
SPIN    1
EXCHANGE 2   0.0  0.0

CONTROL  1      1      1      1      1      1

RPATH    7.29
PRINT    5 0 0 0 0 3

CRITERIA 0.0  0.0  curved  plane
DEBYE    150   176  temp    debye-temp
XANES

POTENTIALS

```

```

*   ipot   z   label
      0   64   Gd
      1   64   Gd

ATOMS
the list of atoms is created by ATOMS program
-----
title    Gd , hcp
! Wycoff, vol.1 p.331
space   hcp
rmax = 9.0
a = 3.6354
c = 5.7817
atom
  Gd   0.33333  0.66667  0.25  center
-----

```

### G.1.3 XMCD sum rules normalization

A procedure for relating XAS to LDOS and XMCD sum rule normalization has been described in a recent paper, A. Nesvizhskii, A. L. Ankudinov and J. J. Rehr, *Normalization and convergence of x-ray absorption sum rules*, Phys. Rev. B (in press) 2001. To carry out this procedure use the appropriate PRINT card option to print 'ratio.dat', which contains  $\rho_0$ ,  $\mu_0$  and their ratio  $\rho_0/\mu_0$ . The calculations have to be performed without the SPIN card. The correction due to the difference between  $j+$  and  $j-$  DOS is already included in the calculated ratio. This correction practically does not affect  $S_z$ , but may increase  $L_z$  by 10 %. No correction is added (or needed) for K and other  $l = 0$  edges.

To get the finite integration range correction you have to run the code 4 times (will automate later): 1) Two times for SPIN  $\pm 1$  to get estimated values of  $S_z$  and  $L_z$  (these are reported on the screen). Ideally one has to adjust amplitudes in subroutine getorb to have approximately correct value for  $S_z$  but the correction is few per cent and does not depend much once you are in the vicinity of correct value; 2) Two times for SPIN  $\pm 1$  and Fermi level shifted to the end of integration range, which can be easily done using EXCHANGE card: e.g EXCHANGE 2 20 0 2 if you stopped integration 20 eV from the edge. The nonzero values for  $S_z$  and  $L_z$  are the corrections.  $dS_z = S'_z - S_z$ . We suggest to take ratio and appropriately scale  $S_z$  and  $L_z$ . This is usually negligible correction for  $L_z$ , but increase  $S_z$  up to 5 per cent.

### G.1.4 SPXAS

For antiferromagnets, the XMCD should be zero. SPXAS is a different technique where you measure spin-up and spin-down signal by measuring intensity of two spin-split Kbeta lines. This corresponds to measuring spin-order relative to the spin on the absorber ( not relative to the external magnetic field as in XMCD). As an example, let us look at the Mn K edge of

antiferromagnetic MnF<sub>2</sub>. Our calculations agree well with experiment in EXAFS region.

Here is the input file for MnF<sub>2</sub>.

```

TITLE  MnF2 (rutile) cassiterite (Wykoff)

HOLE 1  1.0      1=k edge, s0^2=1.0
SPIN  -2

CONTROL  1      1      1      1      1      1
PRINT    0      0      0      2      0      4
EXCHANGE 0  0.0  0.0
CORRECTIONS 0.0  0.0

RPATH  10.0
XANES

PCRITERIA      0.8      40.0
*CRITERIA      curved  plane
CRITERIA      0.0      0.0
*DEBYE         temp    debye-temp
  DEBYE        300      350
NLEG           4

POTENTIALS
*  ipot  z  label
    0  25  Mnup
    1   9   F
    2  25  Mnup
    3  25  Mndown

ATOMS
  0.0000  0.0000  0.0000  0  Mnup  0.0000
  1.4864  1.4864  0.0000  1  F      2.1021
 -1.4864 -1.4864  0.0000  1  F      2.1021
  0.9503 -0.9503  1.6550  1  F      2.1319
  0.9503 -0.9503 -1.6550  1  F      2.1319
 -0.9503  0.9503  1.6550  1  F      2.1319
 -0.9503  0.9503 -1.6550  1  F      2.1319
  0.0000  0.0000 -3.3099  3  Mndown 3.3099
  0.0000  0.0000  3.3099  3  Mndown 3.3099
 -3.3870  1.4864  0.0000  1  F      3.6988
  3.3870 -1.4864  0.0000  1  F      3.6988
  1.4864 -3.3870  0.0000  1  F      3.6988
 -1.4864  3.3870  0.0000  1  F      3.6988

```



---

2.4367	2.4367	-1.6550	2	Mnup	3.8228
-2.4367	-2.4367	-1.6550	2	Mnup	3.8228
2.4367	-2.4367	-1.6550	3	Mndown	3.8228
-2.4367	-2.4367	1.6550	3	Mndown	3.8228
-2.4367	2.4367	-1.6550	3	Mndown	3.8228
2.4367	2.4367	1.6550	3	Mndown	3.8228
2.4367	-2.4367	1.6550	2	Mnup	3.8228
-2.4367	2.4367	1.6550	2	Mnup	3.8228
...					

END

## G.2 XNCD

XNCD calculations due in nonmagnetic materials due to E1-E2 mechanism is performed, i.e. due to cross electric dipole -quadrupole transitions. It has to be used with XANES card. We performed calculations for  $\text{LiIO}_3$  and found results very similar to previous multiple scattering XNCD calculations.

## G.3 X-ray elastic scattering amplitude $f'$ , DAFS, DANES

All necessary components to obtain elastic scattering amplitude can be calculated with FEFF8.10. Thus Thomson scattering amplitudes are written in file 'fpf0.dat', elastic amplitude near some specific edge are calculated with DANES card while far from the edge with FPRIME card which neglects solid state effects on  $f'$ . The  $f''$  can be obtained with XANES card. The formula connecting  $f''$  and absorption cross section  $\sigma$  is (in atomic units)  $f'' = \omega c \sigma / 4 / \pi$ .